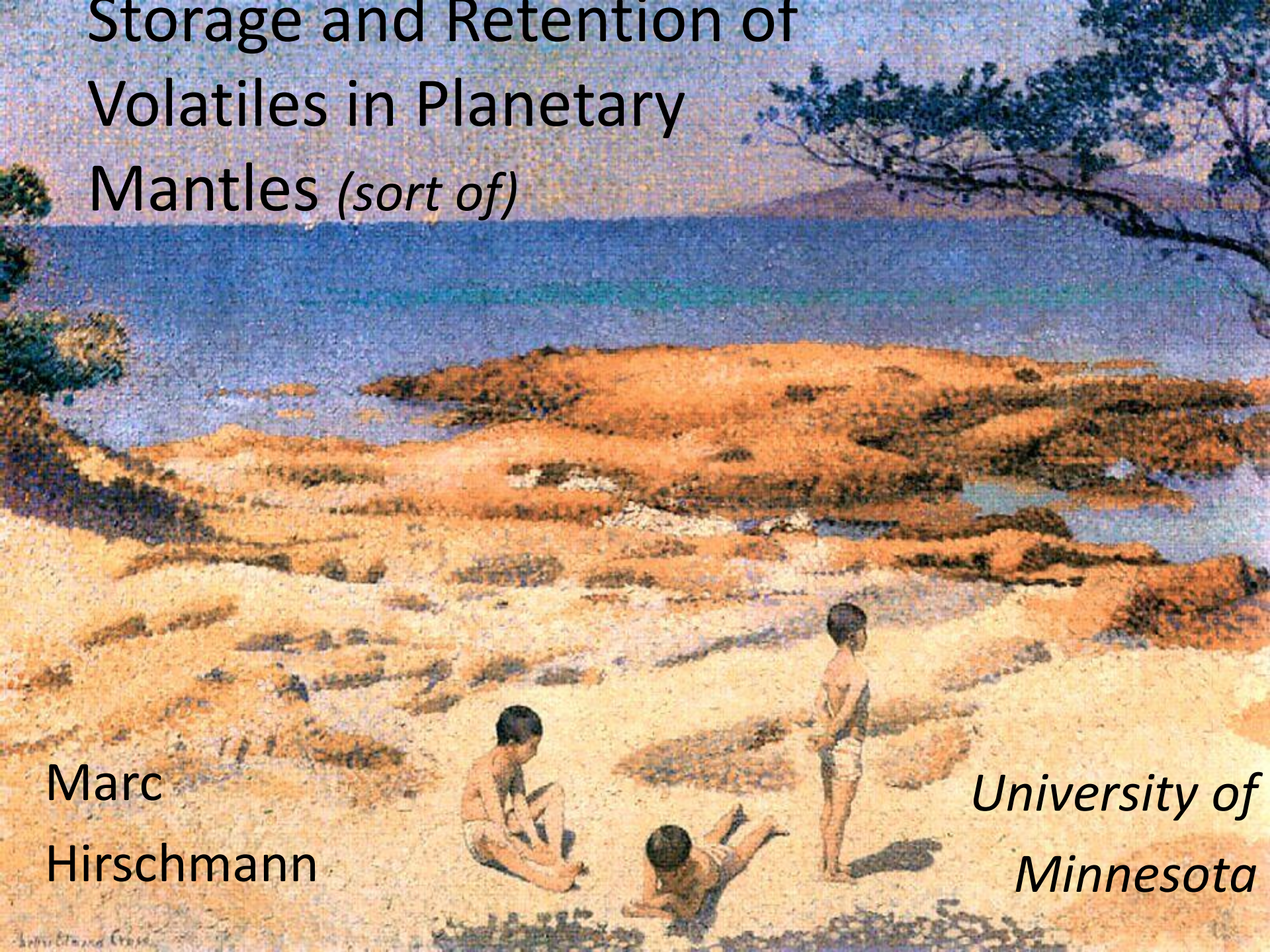


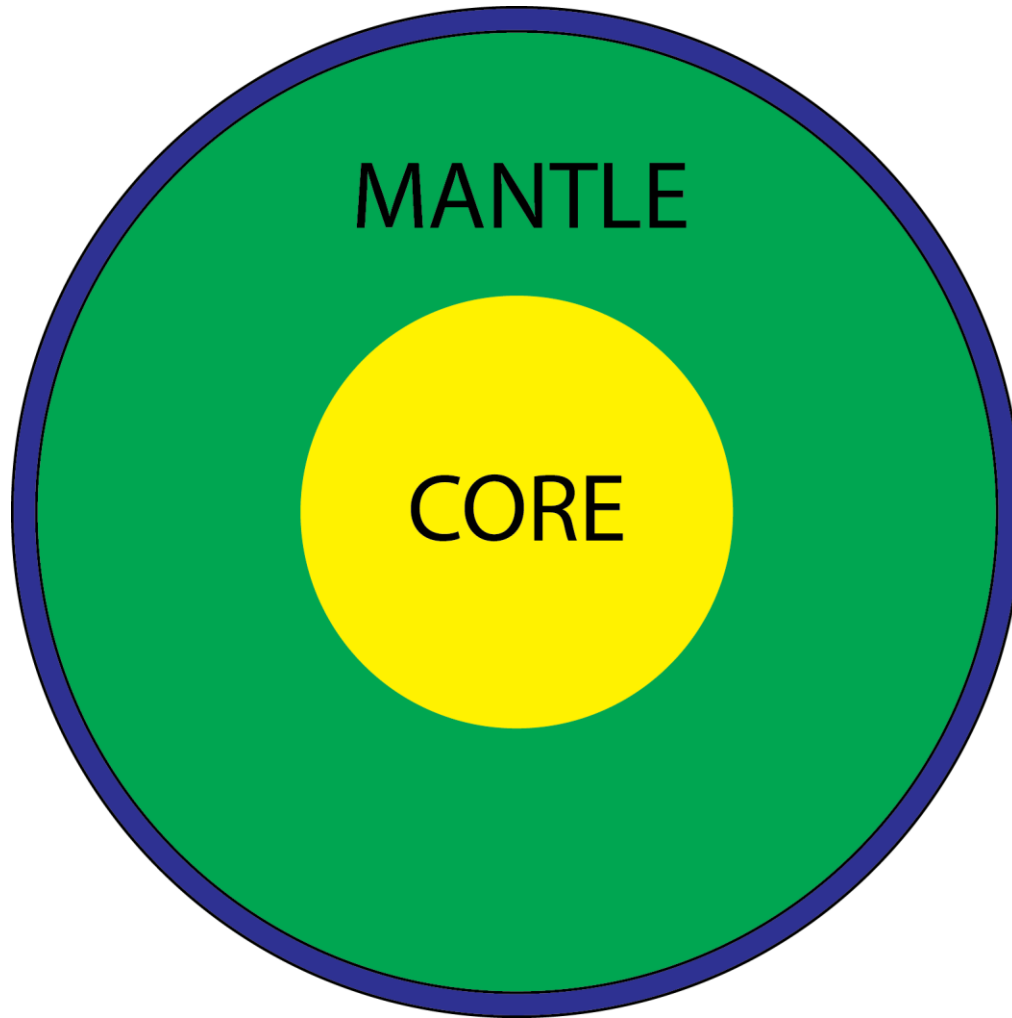
Storage and Retention of Volatiles in Planetary Mantles *(sort of)*

Marc
Hirschmann

*University of
Minnesota*

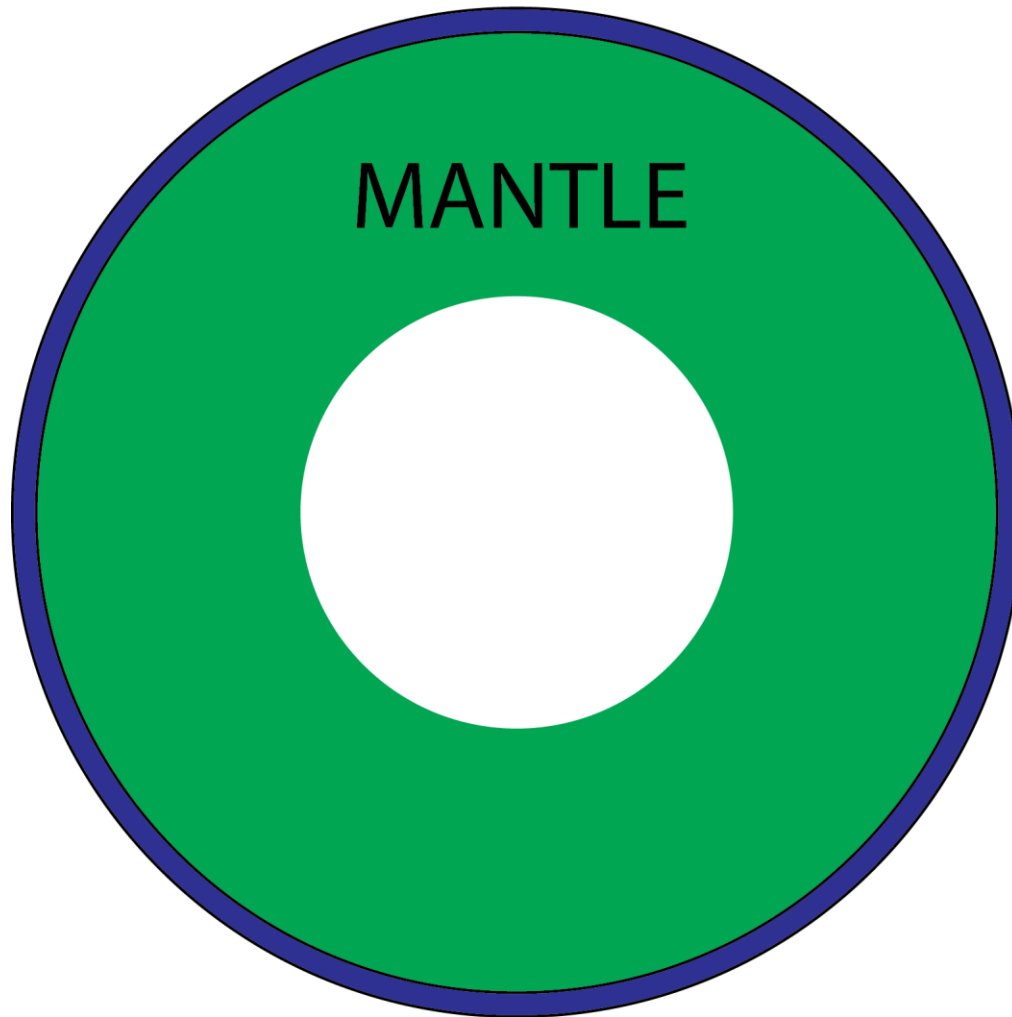


- What is the inventory of volatiles (H,C,N,S) in the bulk silicate Earth?
- How do these compare to potential cosmochemical sources?
- How does this constrain processes of volatile acquisition during accretion, differentiation, loss?



Exosphere
(=fluid envelopes,
sediments,crust)

“Bulk Silicate Earth” (BSE)

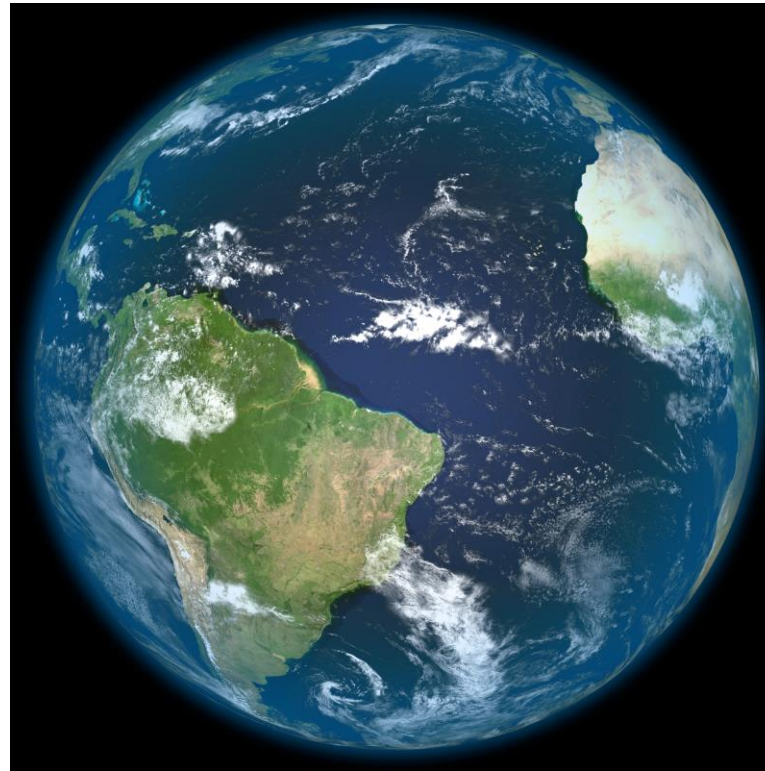


Exosphere
(=fluid envelopes,
sediments, crust)

Hydrogen in the Exosphere (Exosphere=everything above the Moho)

(Lecuyer et al. 1998)

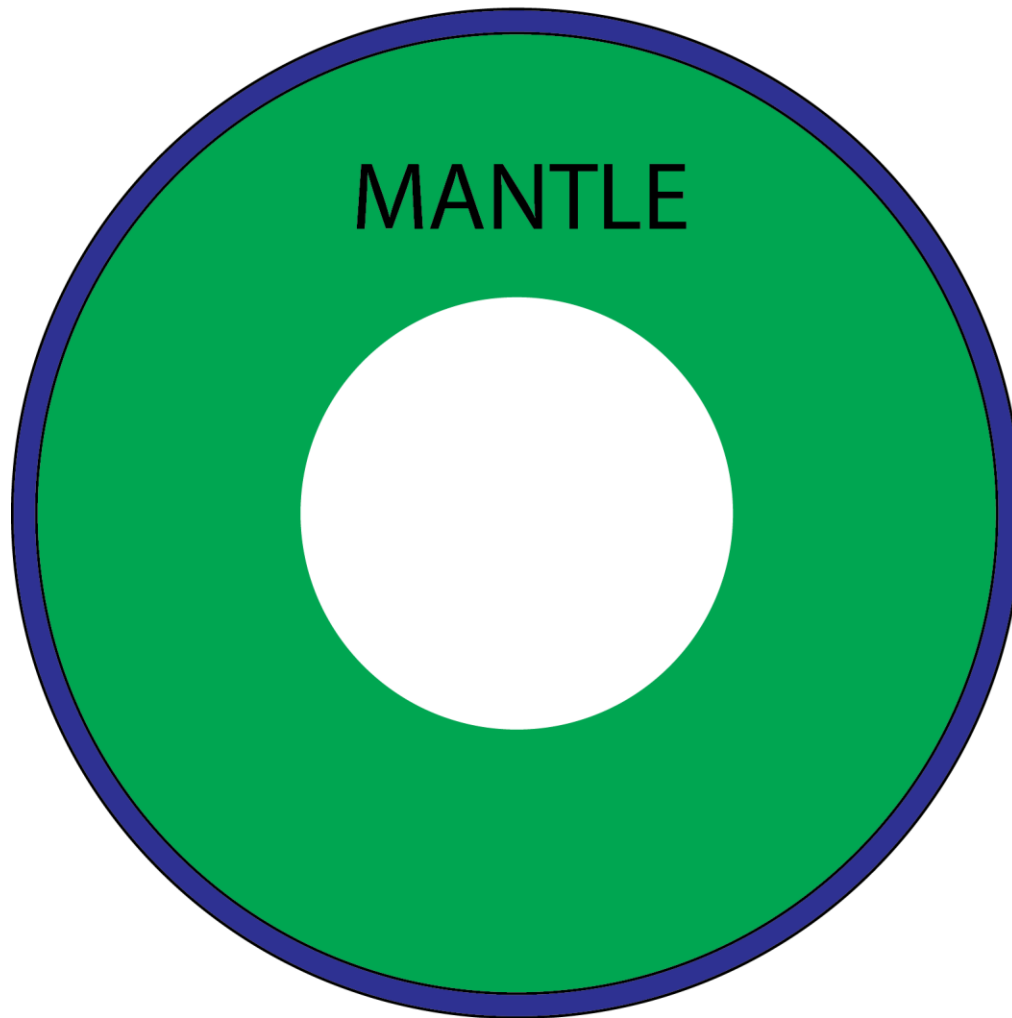
	<u>Grams H₂O</u>	<u>Grams H</u>
Oceans	1.4×10^{24}	1.56×10^{23}
Other	2×10^{23}	2.22×10^{22}
Total	1.6×10^{24}	1.78×10^{23}



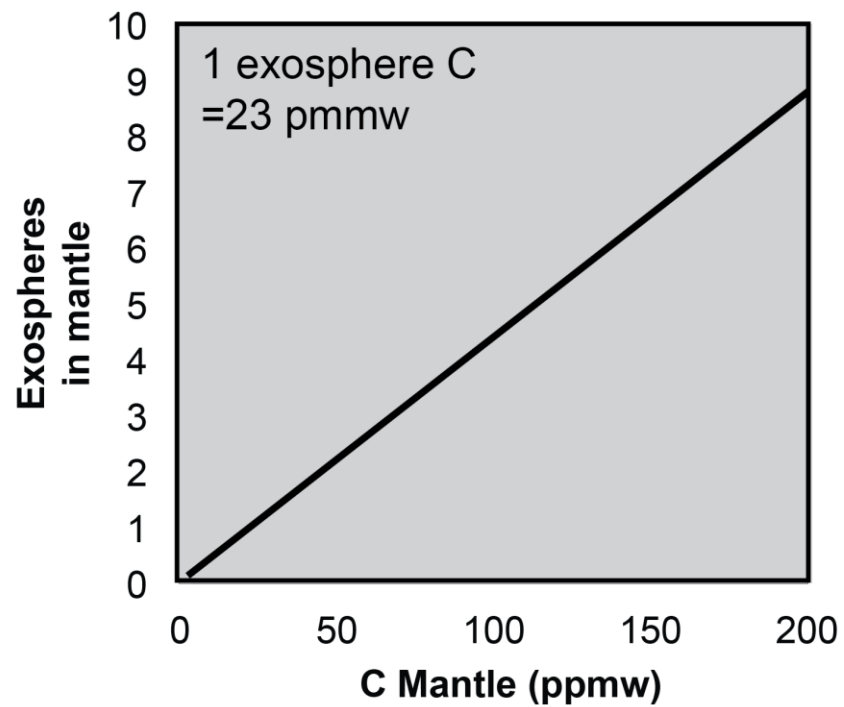
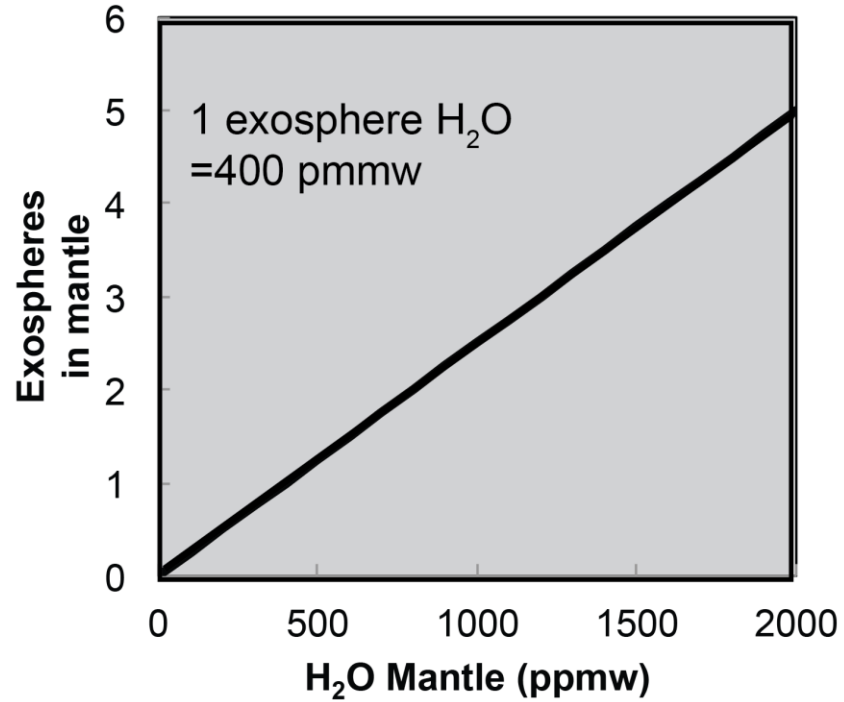
Exosphere Carbon	Moles CO₂	Grams C
Sleep&Zahnle '02		
Sediments	5.88×10^{21}	7.06×10^{22}
Oceans	3.31×10^{18}	3.97×10^{19}
Oceanic Crust	1.20×10^{21}	1.44×10^{22}
Total	7.08×10^{21}	8.50×10^{22}
Hayes&Waldbauer '06	8.50×10^{21}	1.02×10^{23}
Holser '89	7.64×10^{21}	9.17×10^{22}
Average		$9.29 \pm 0.86 \times 10^{22}$



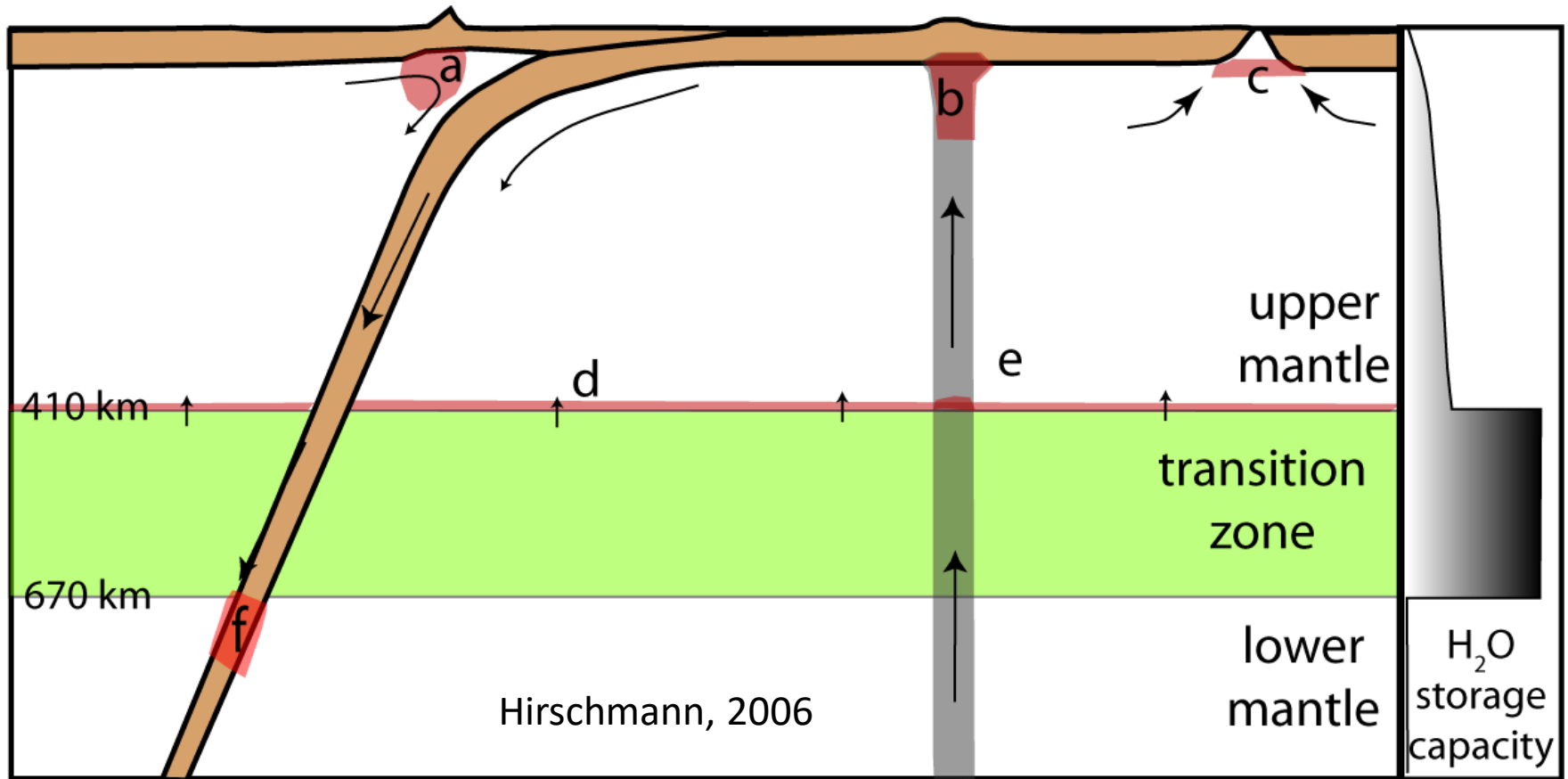
“Bulk Silicate Earth” (BSE)

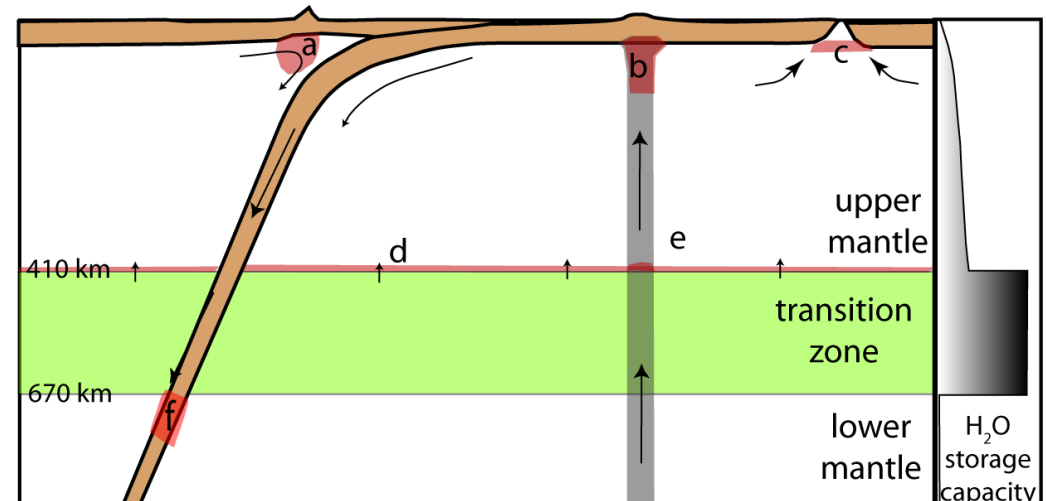
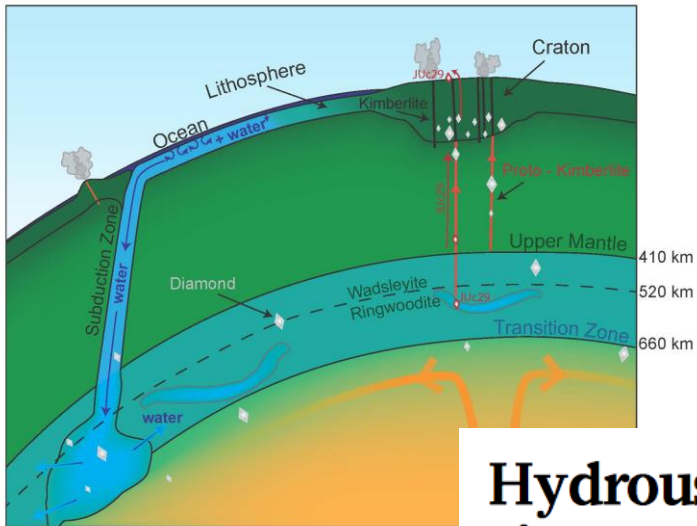
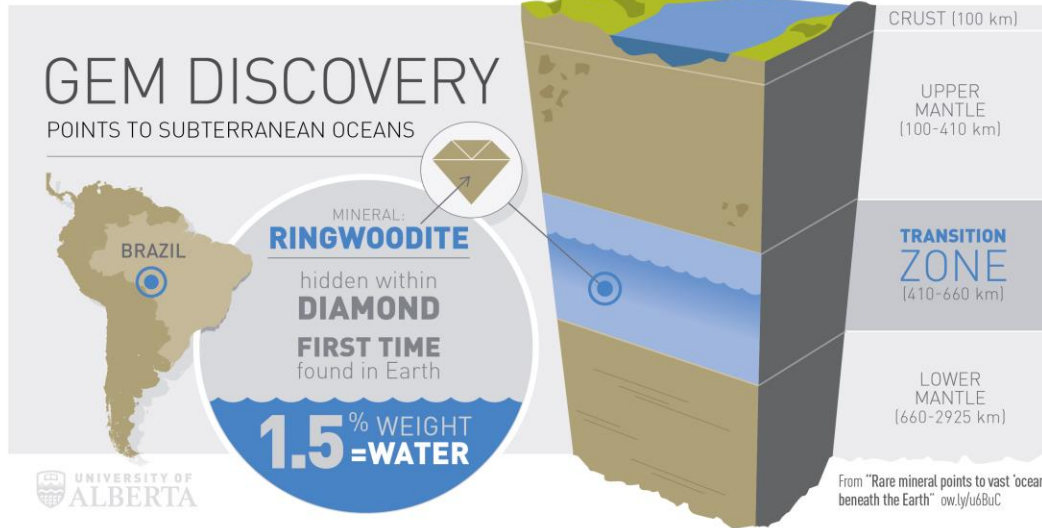


Exosphere
(=fluid envelopes,
sediments, crust)



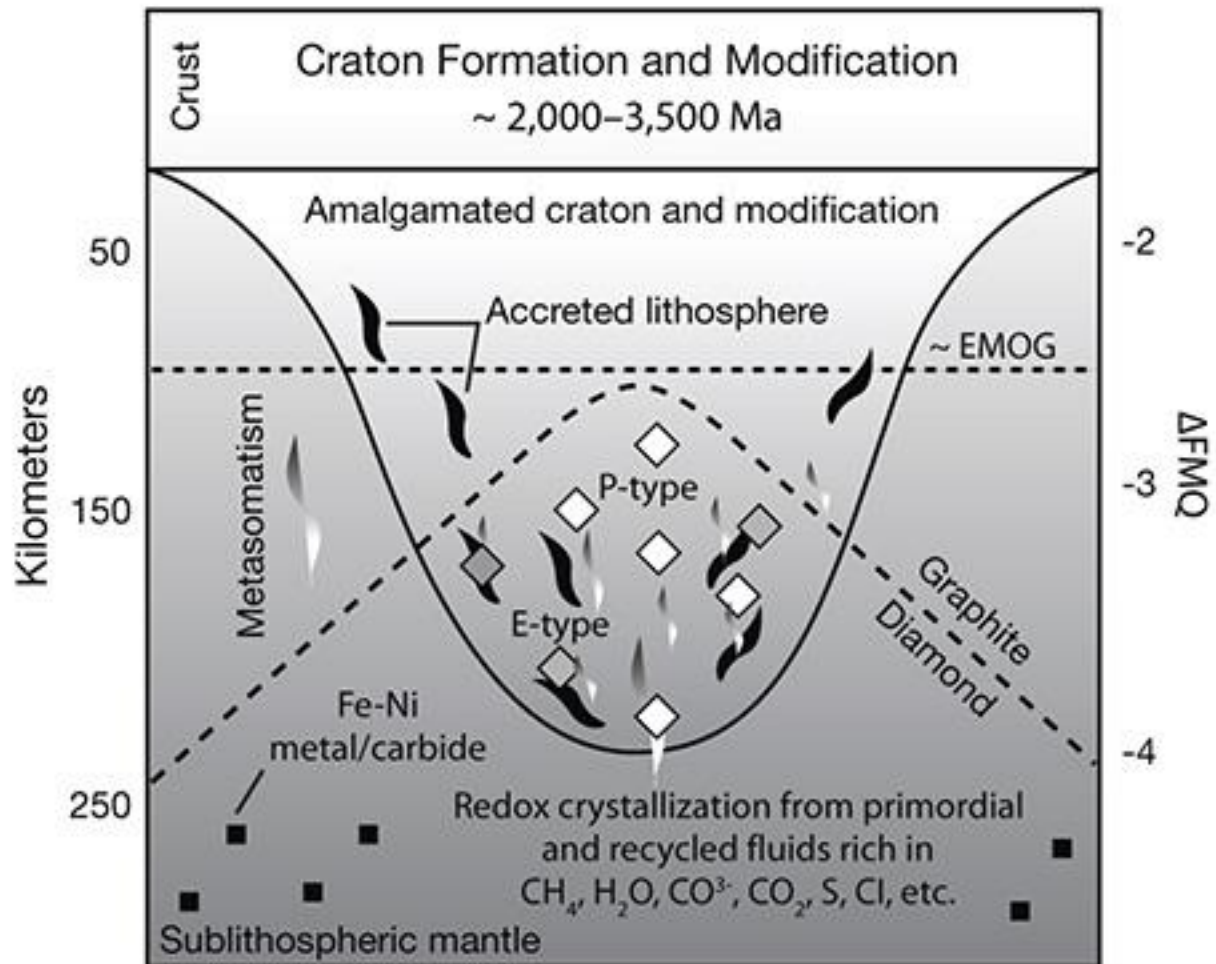
Schematic Features of Water Storage (and hydrous melting) the Mantle



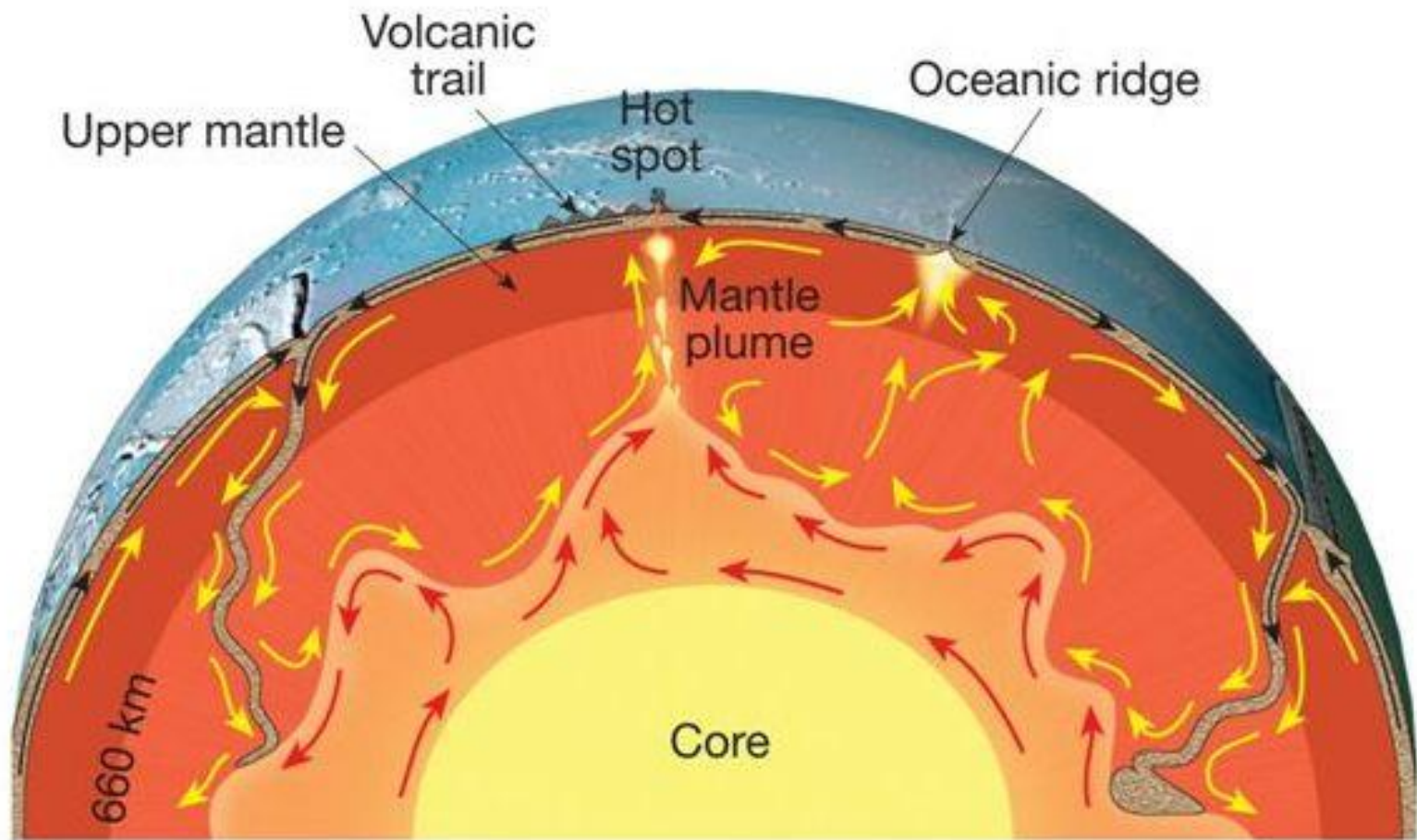


Hydrous mantle transition zone indicated by ringwoodite included within diamond

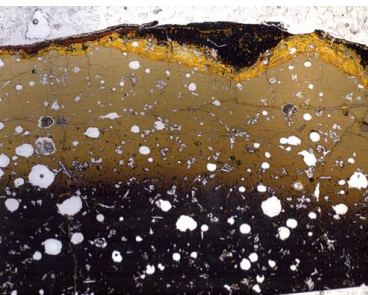
D. G. Pearson¹, F. E. Brenker², F. Nestola³, J. McNeill⁴, L. Nasdala⁵, M. T. Hutchison⁶, S. Matveev¹, K. Mather⁴, G. Silversmit⁷, S. Schmitz², B. Vekemans⁷ & L. Vincze⁷



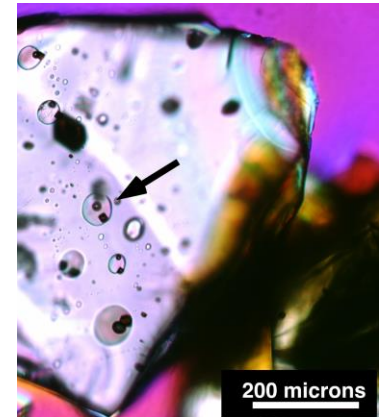
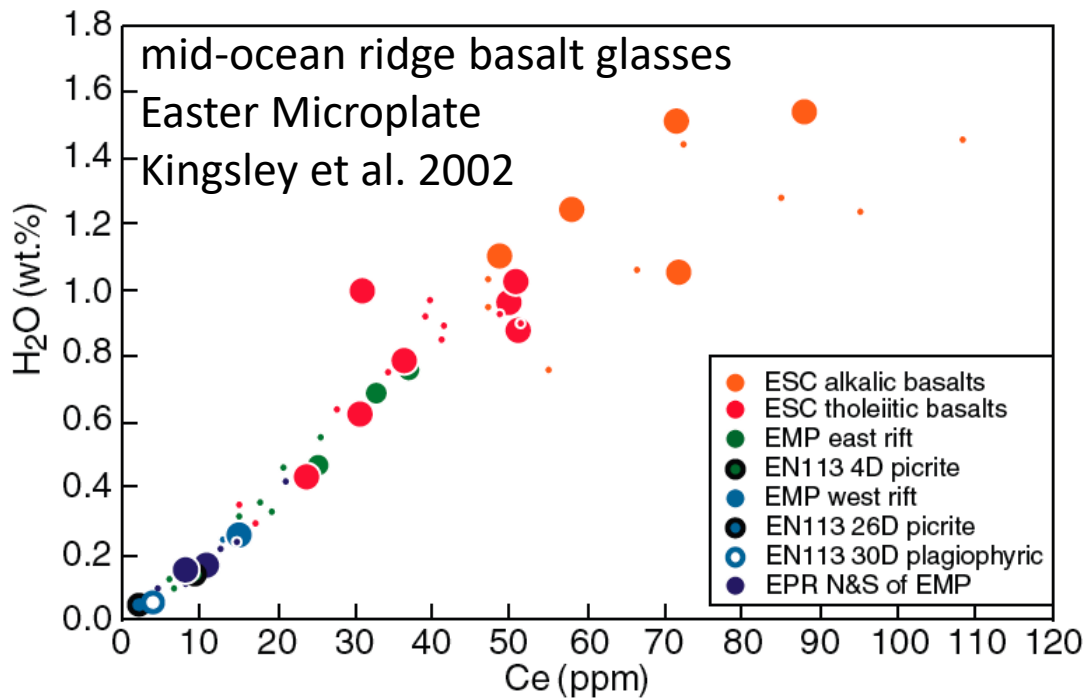
Shirey, 2013



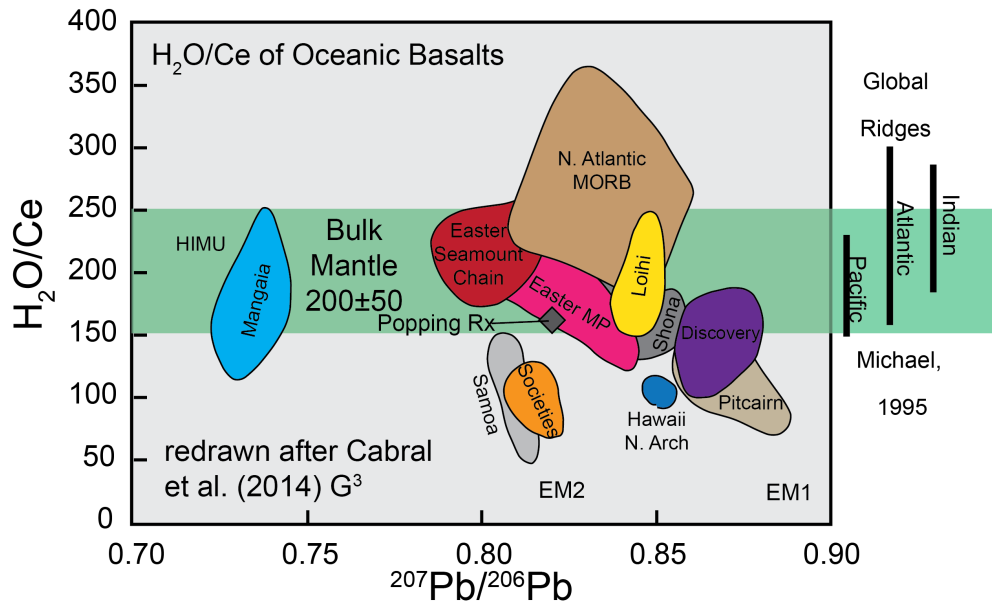
C.

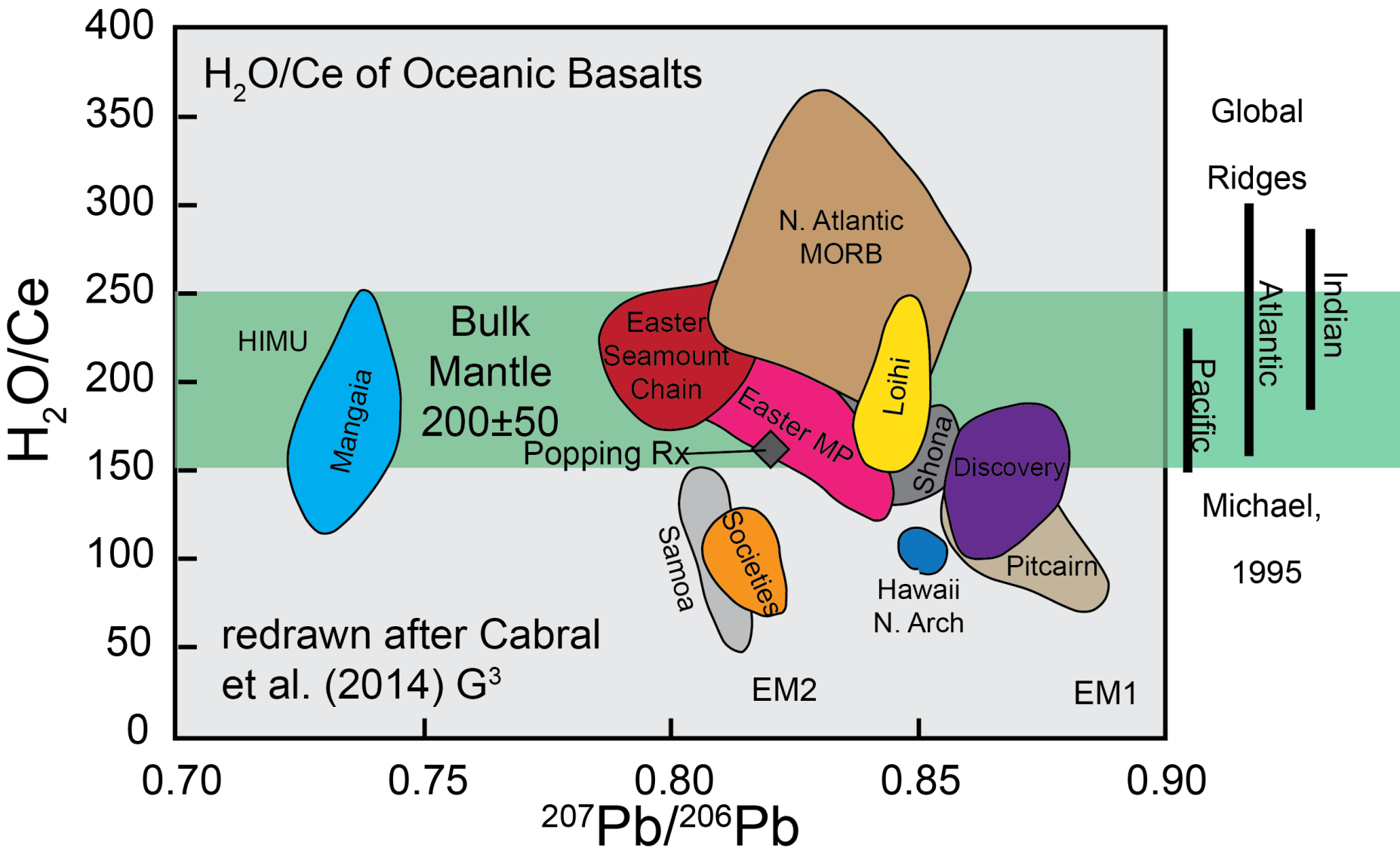


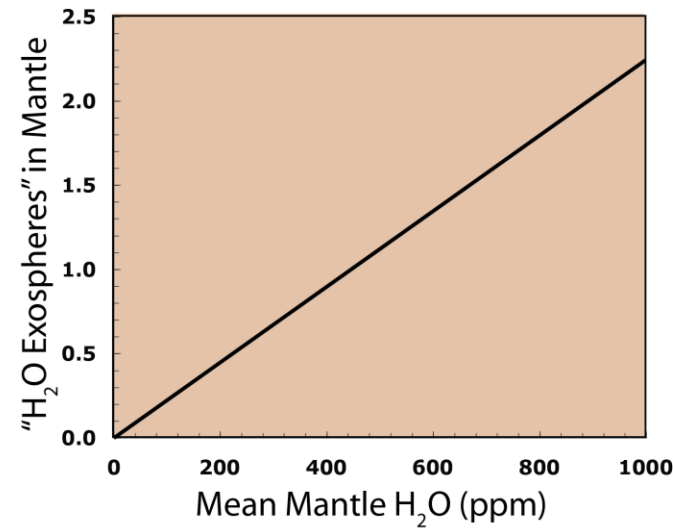
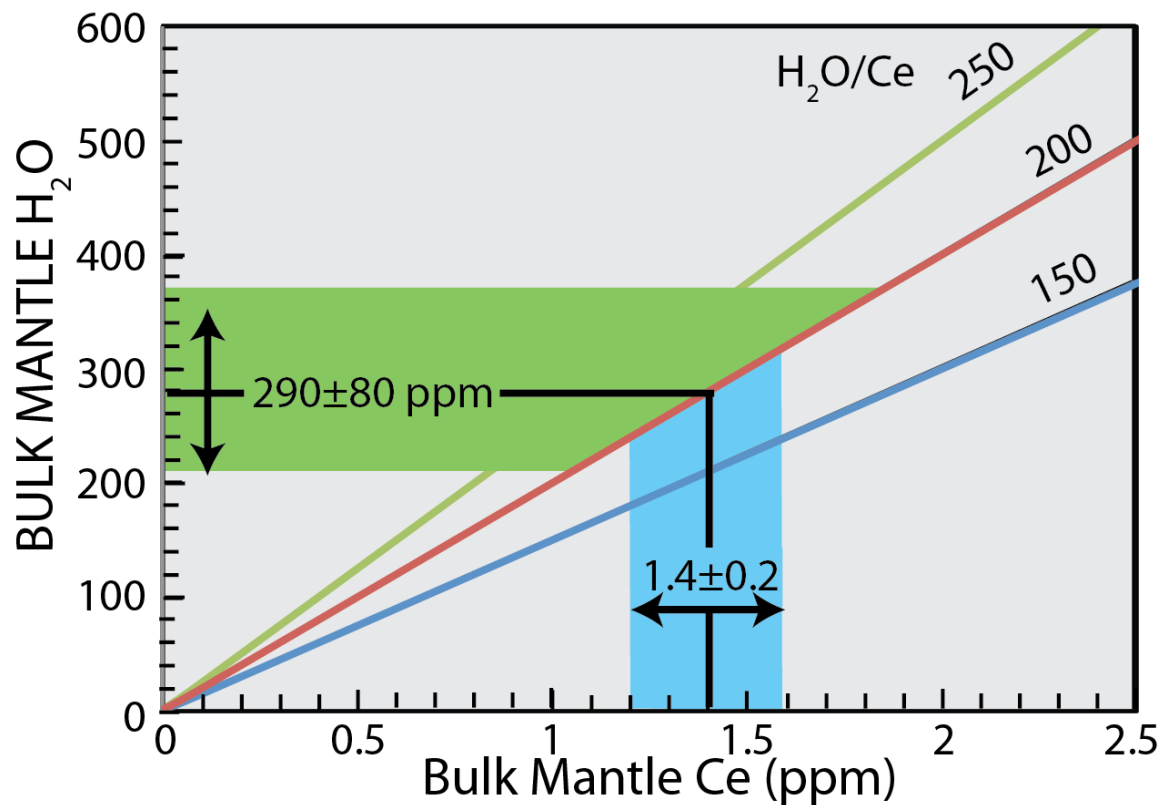
Mid-Ocean Ridge
Quenched Pillow Glass



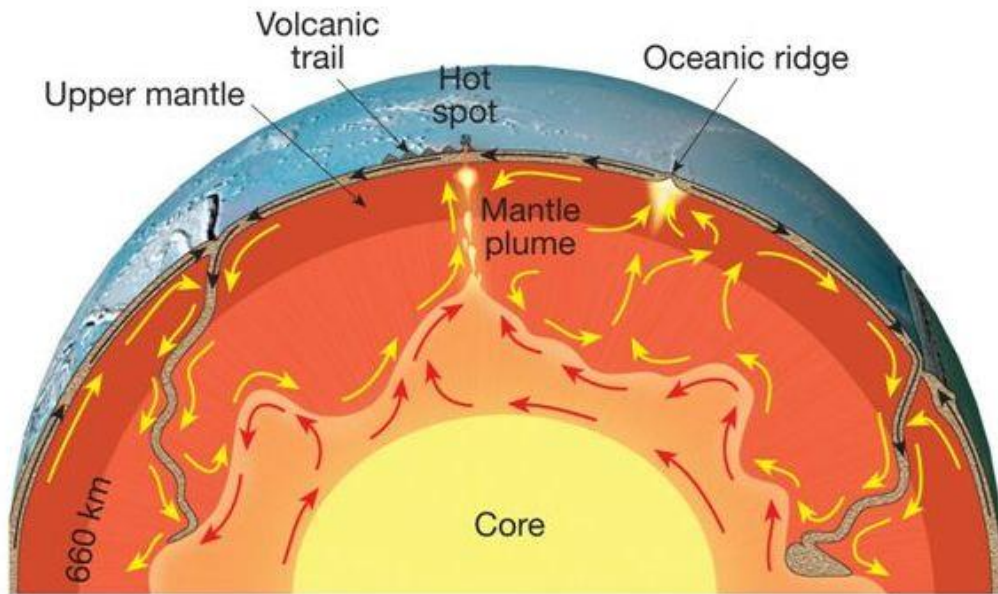
Olivine-hosted
Melt Inclusions
(in MORB or OIB)





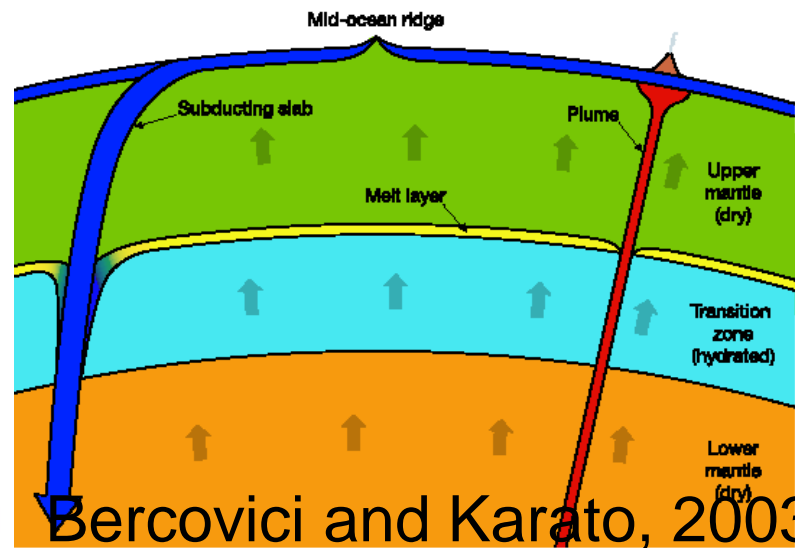


[McDonough&Sun, 1995, minus Continental Crust]

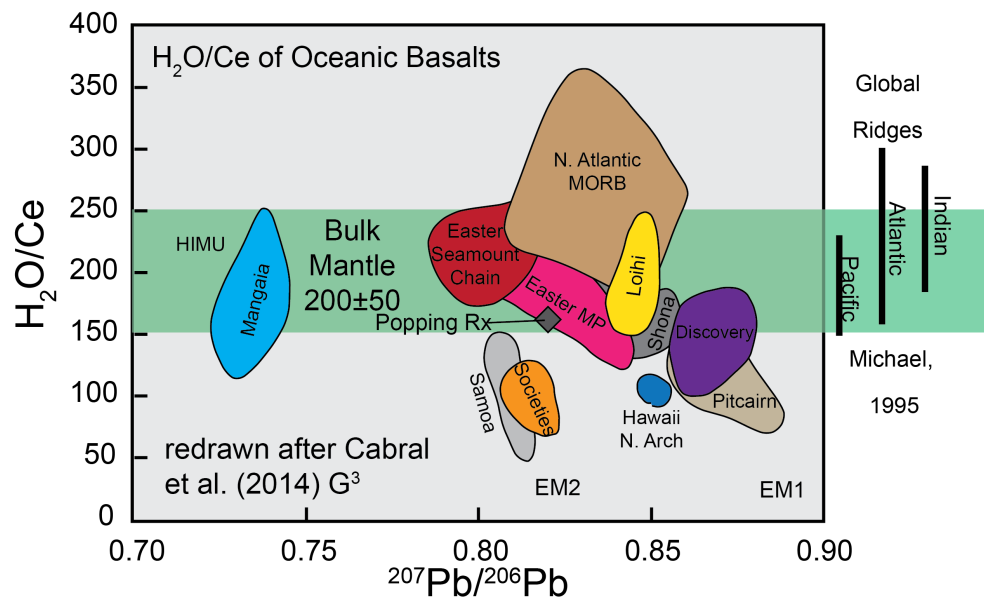
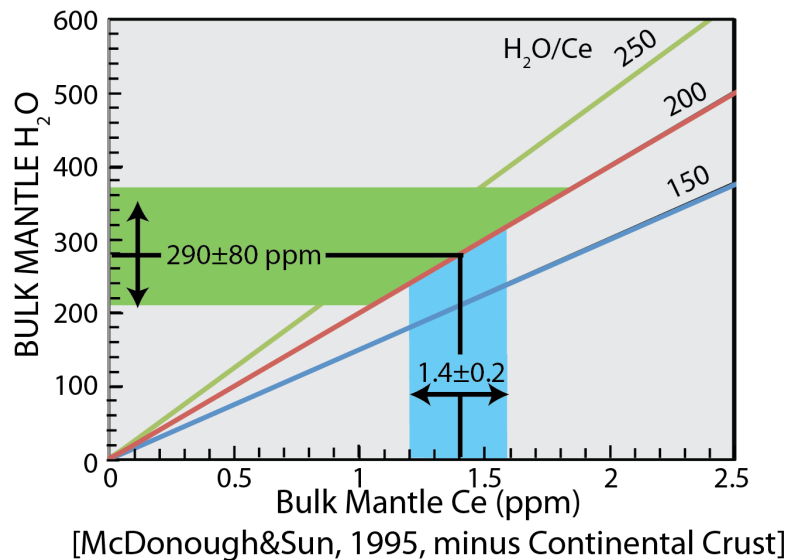


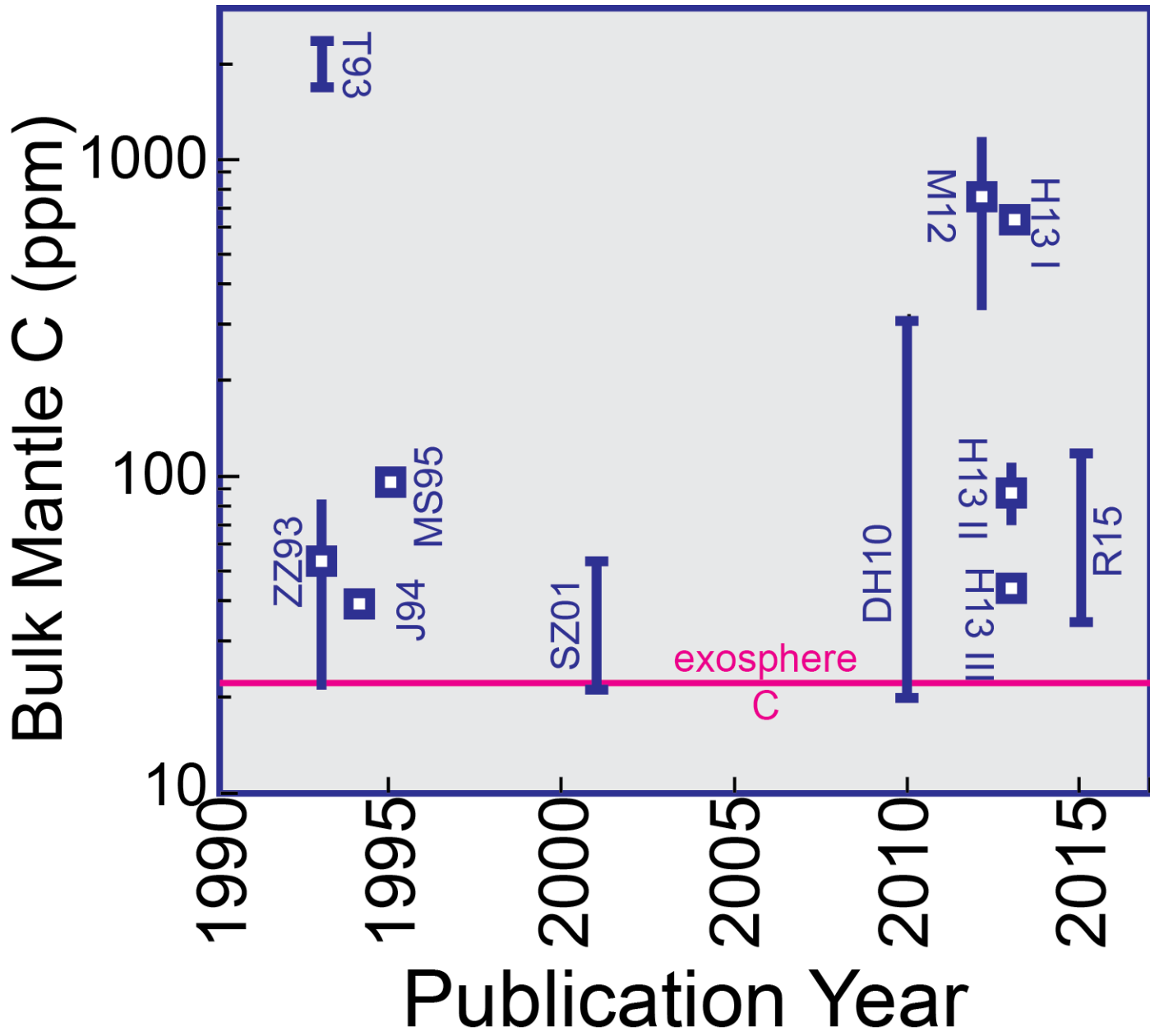
C.

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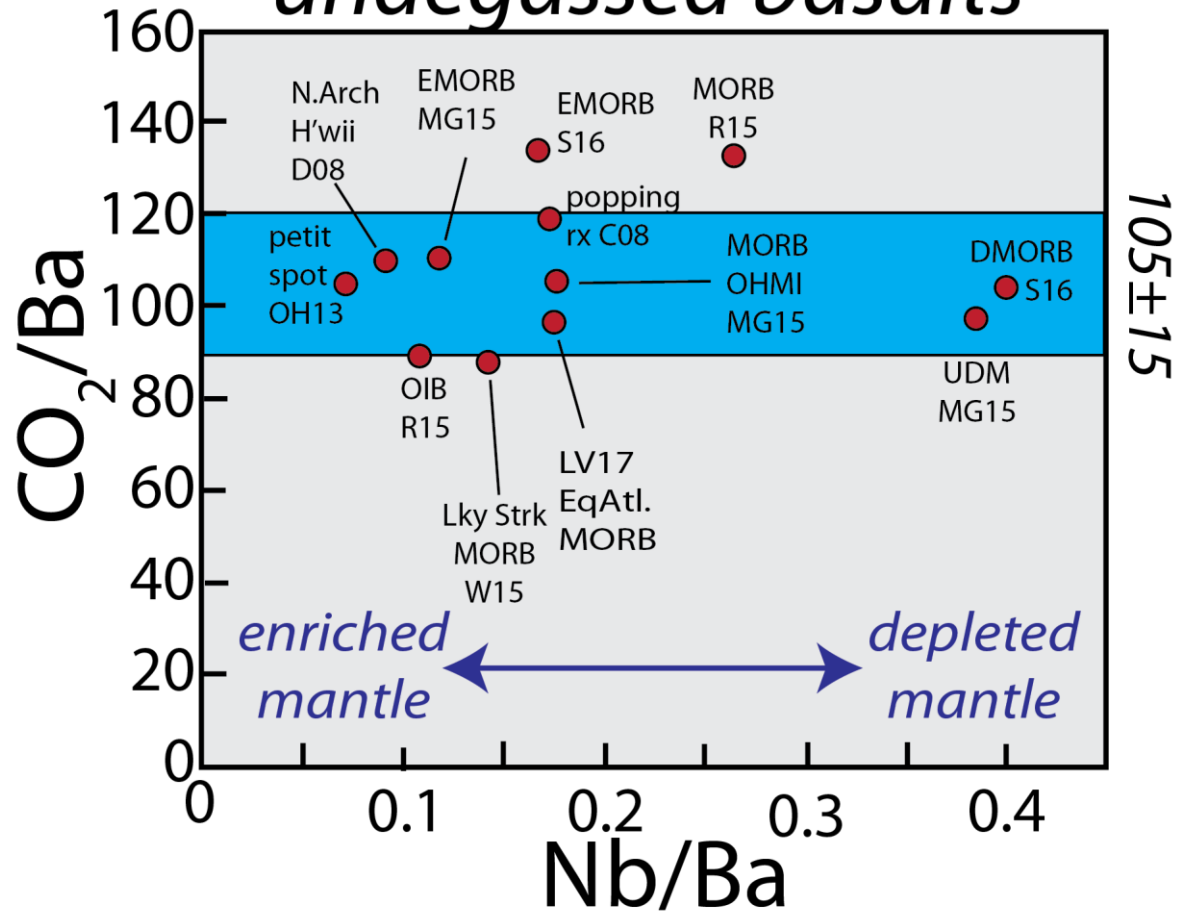


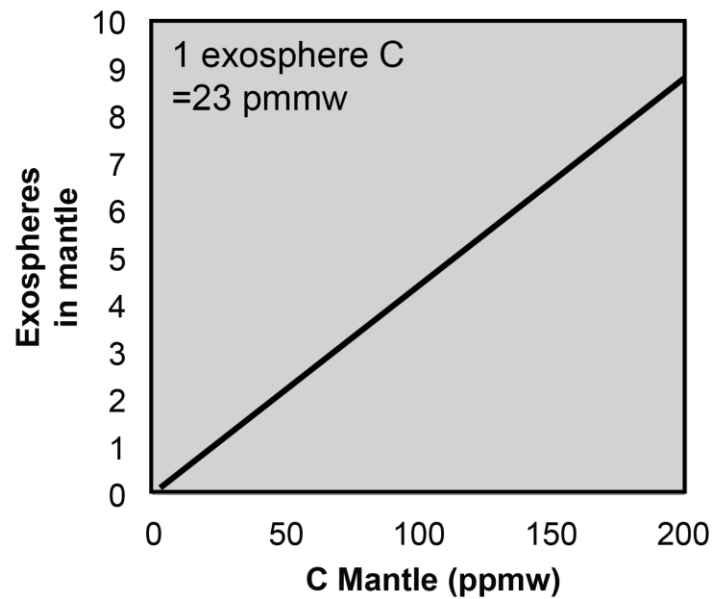
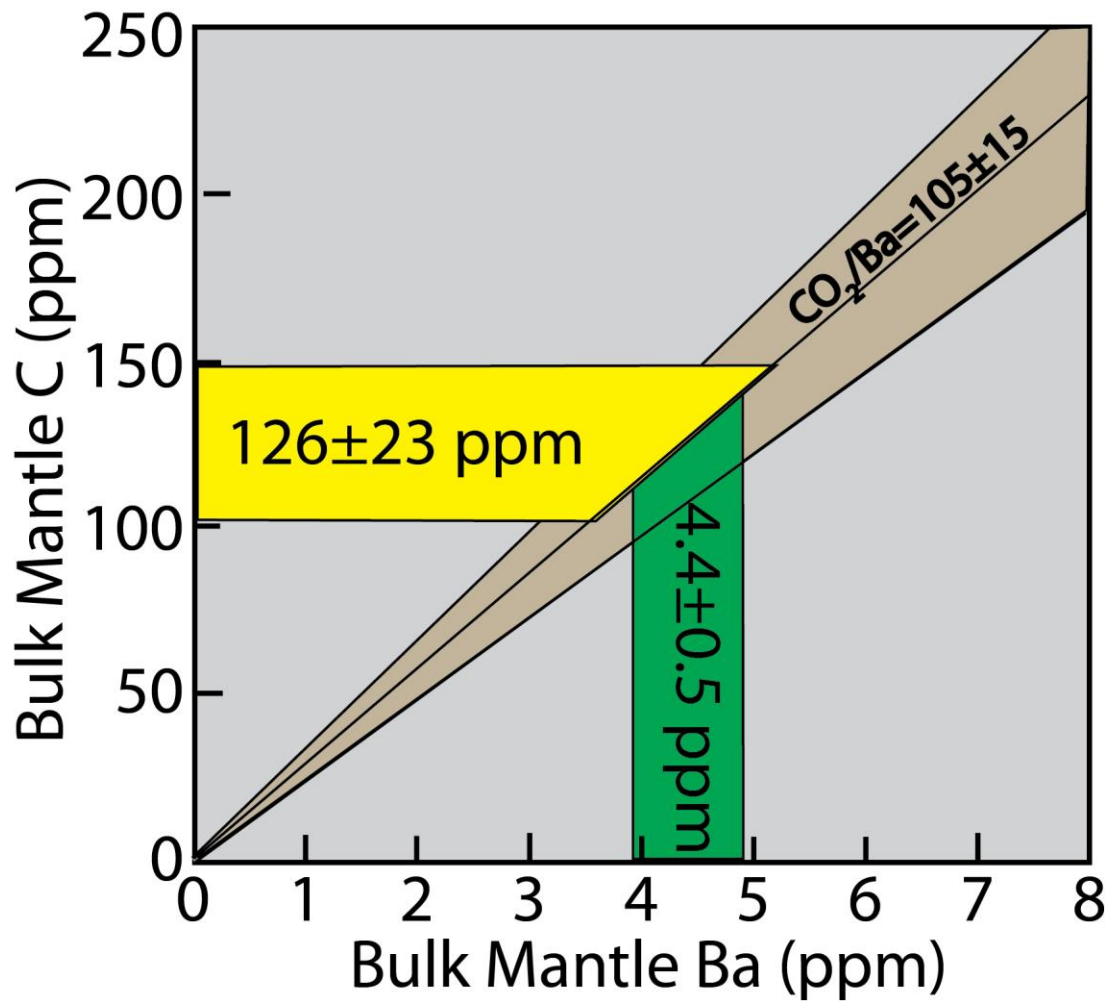
Bercovici and Karato, 2003





CO_2/Ba of undegassed basalts



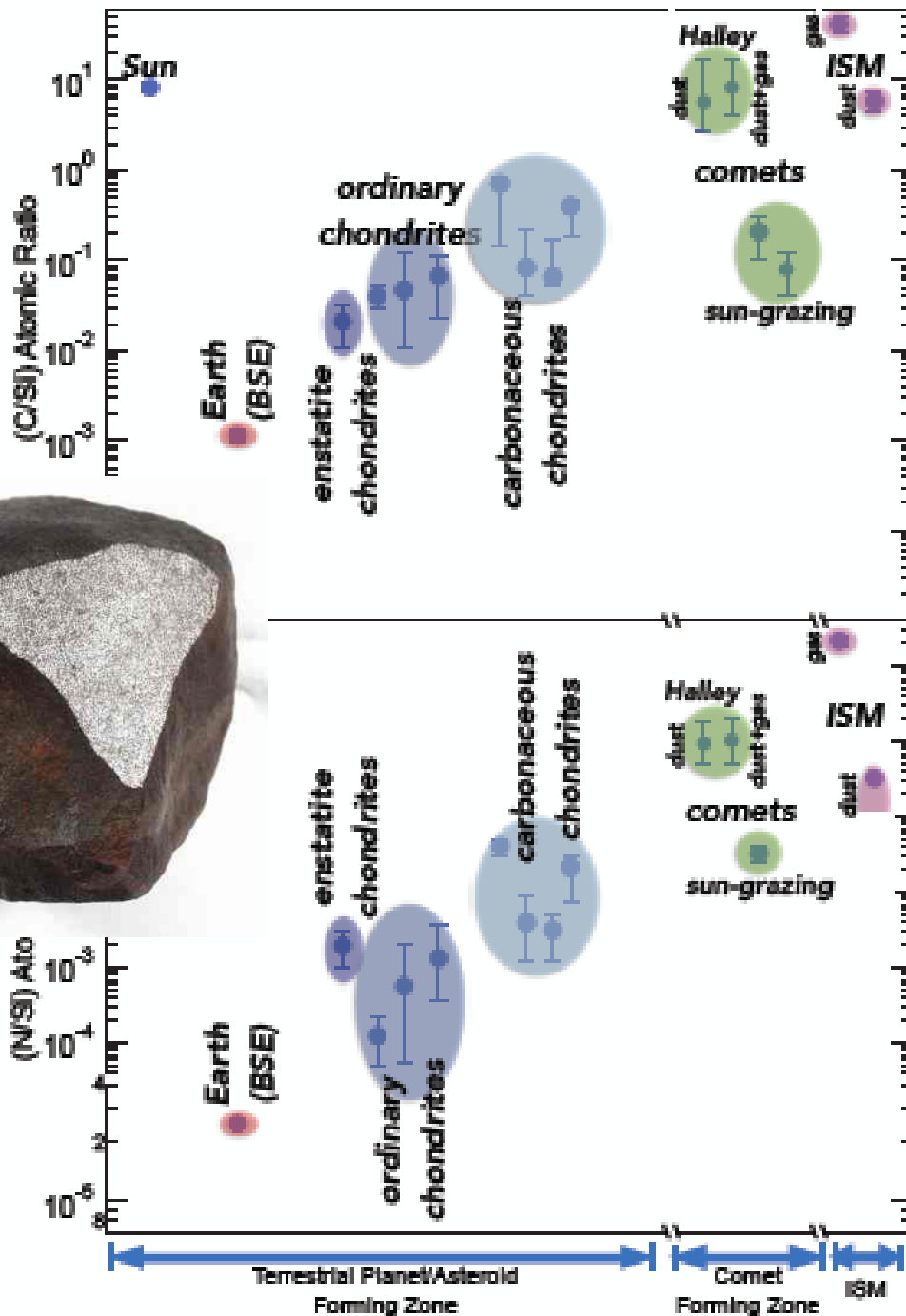




Enstatite chondrite

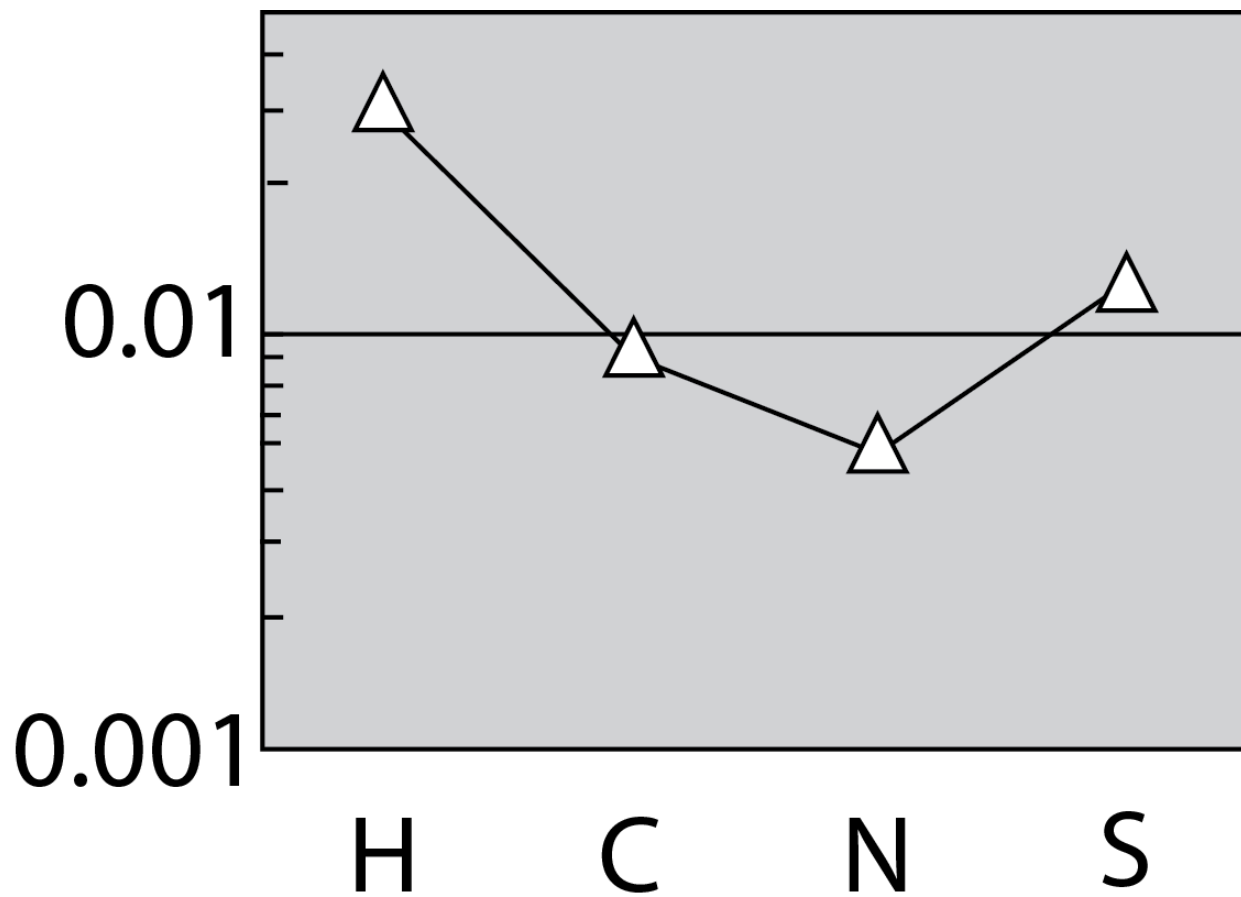


carbonaceous chondrite

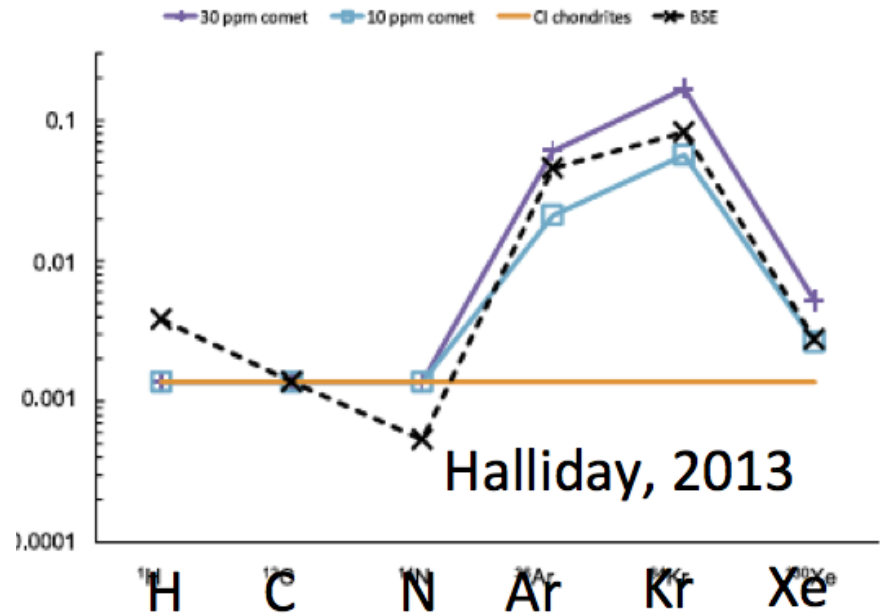
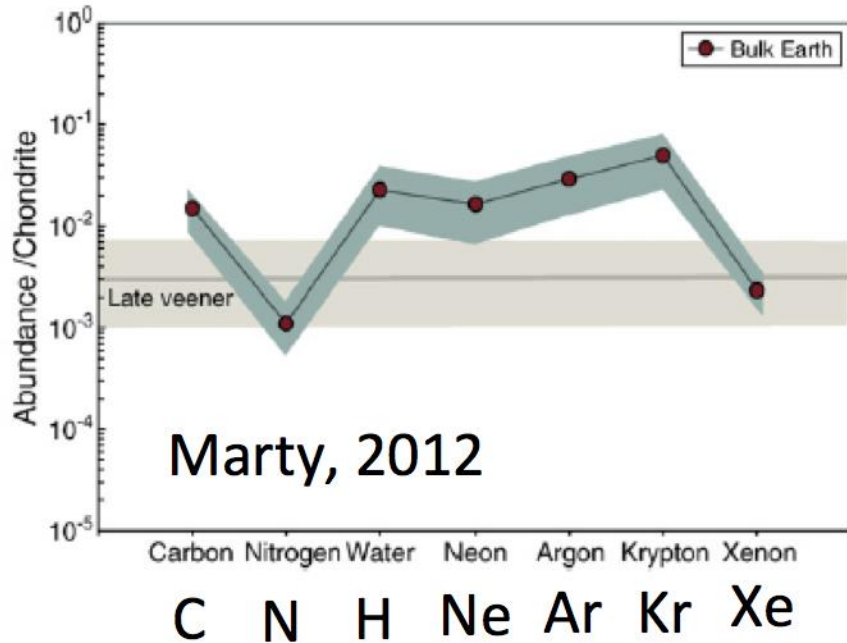


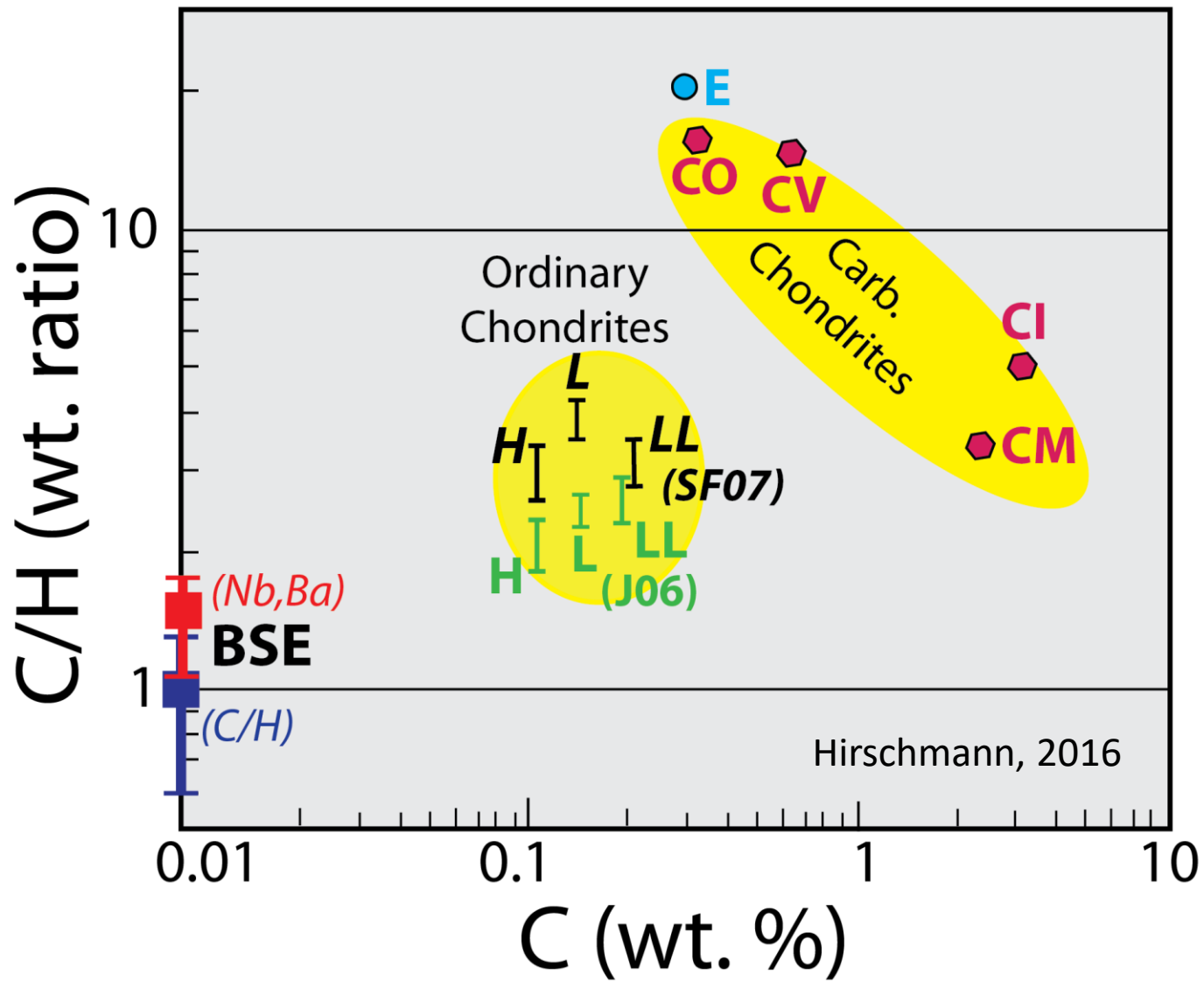
Bergin et al. 2015
PNAS

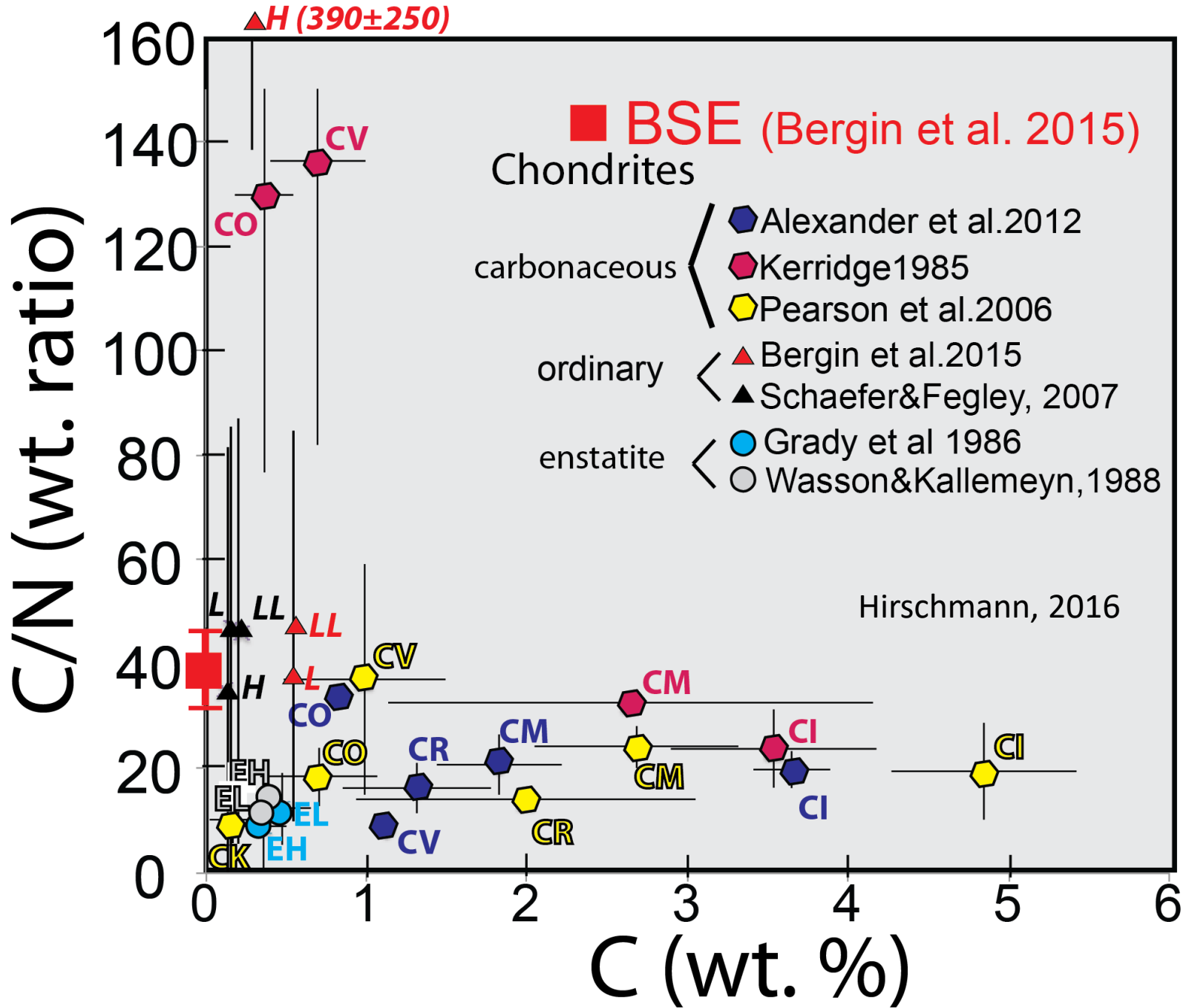
Bulk Silicate Earth
/Chondrite (CI)

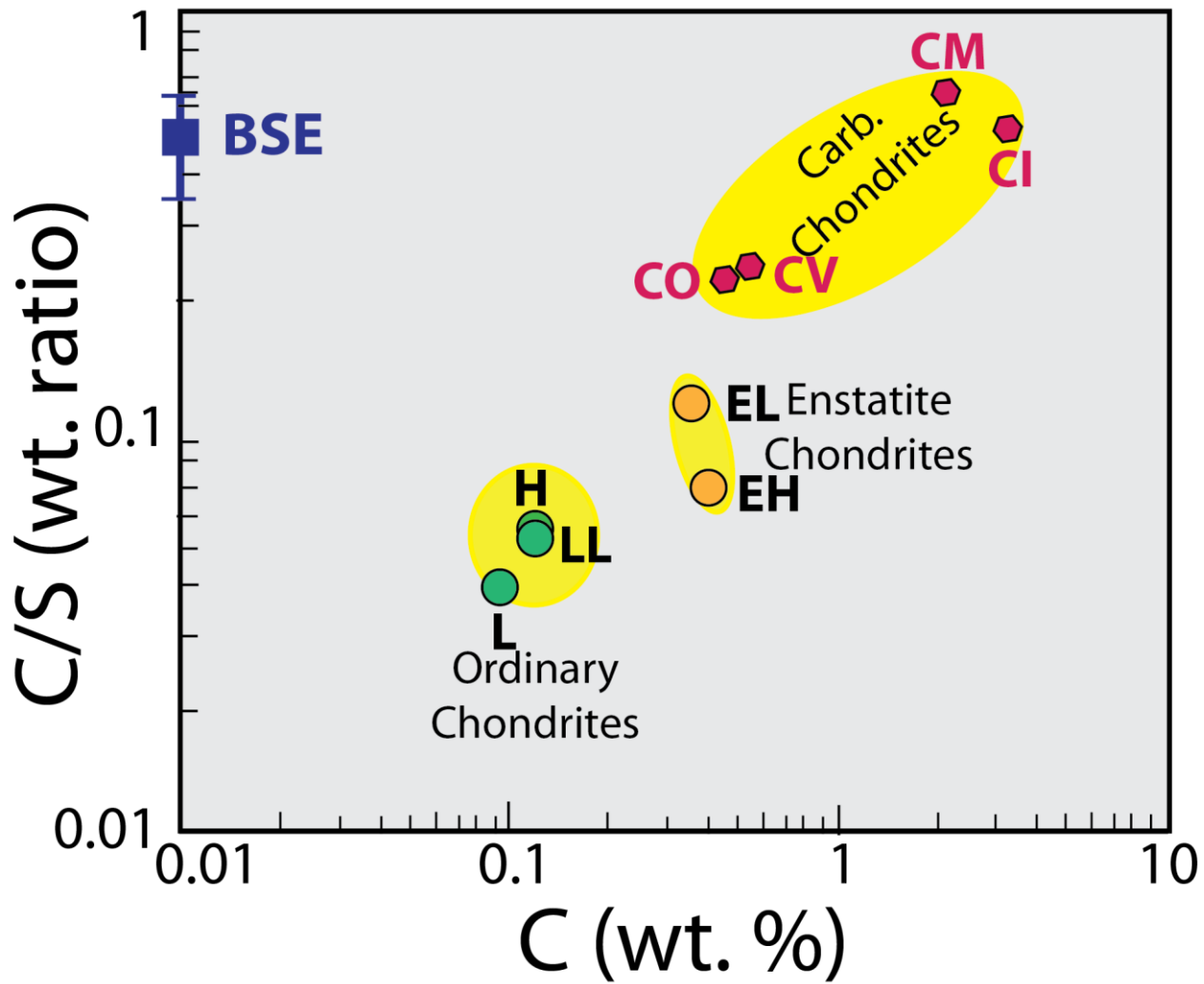


The Bulk Silicate Earth Abundances of Major Volatiles Are FRACTIONATED compared to chondritic abundances

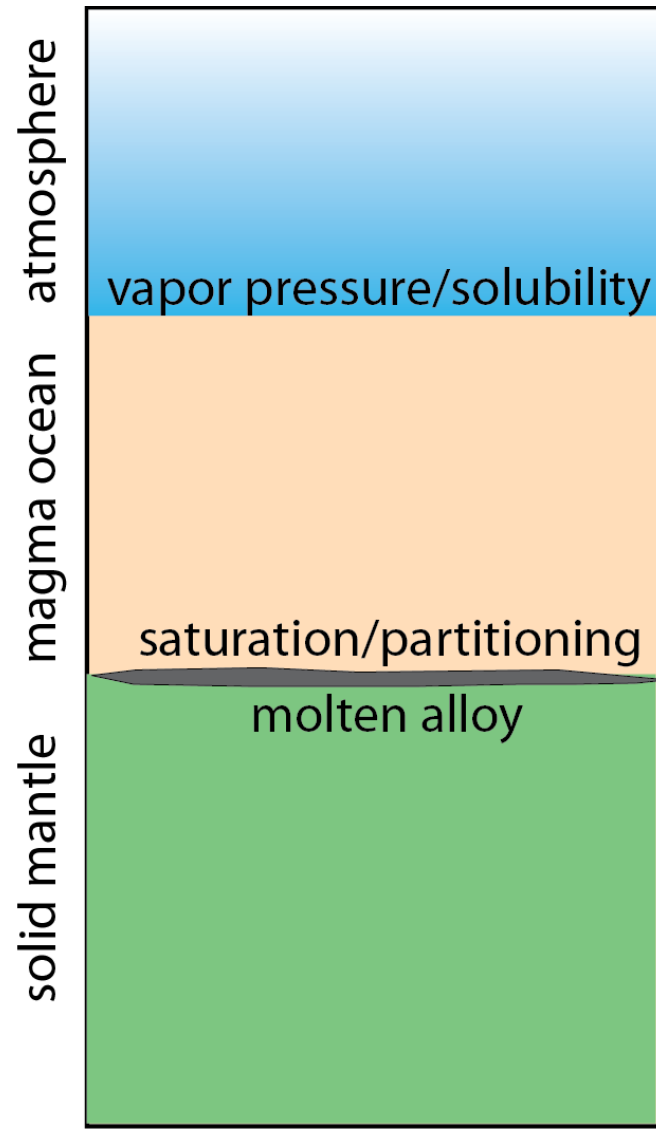


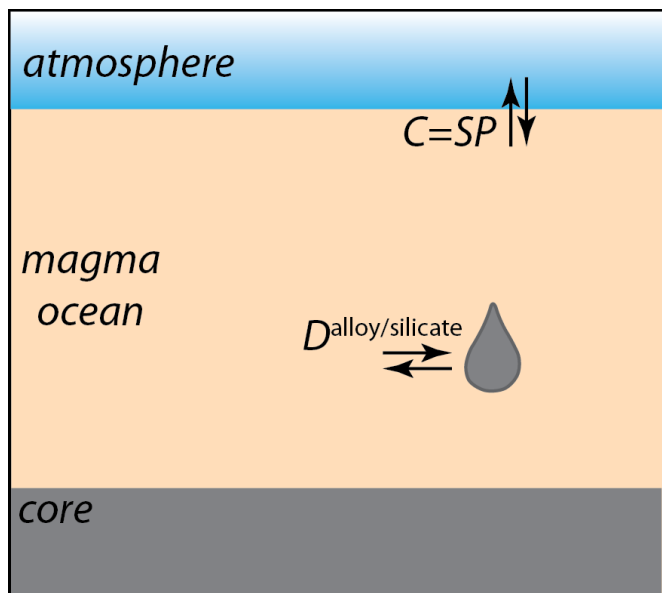






Consider a magma ocean: It produces (at least) 3 reservoirs
silicate, core, and atmosphere

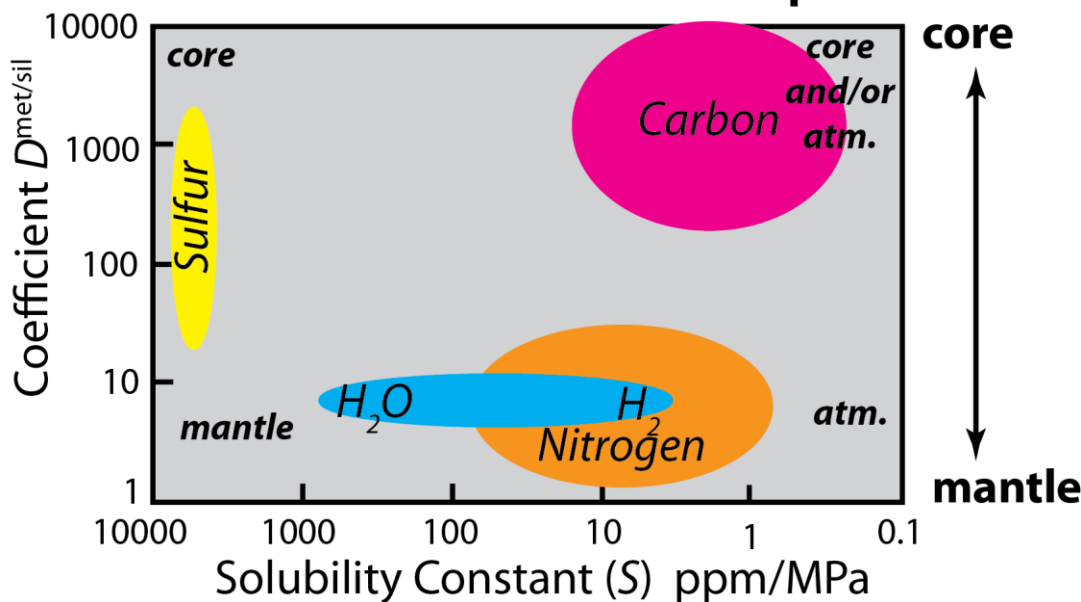




Hirschmann, 2016

Chief Reservoirs in a Largely Molten Planet

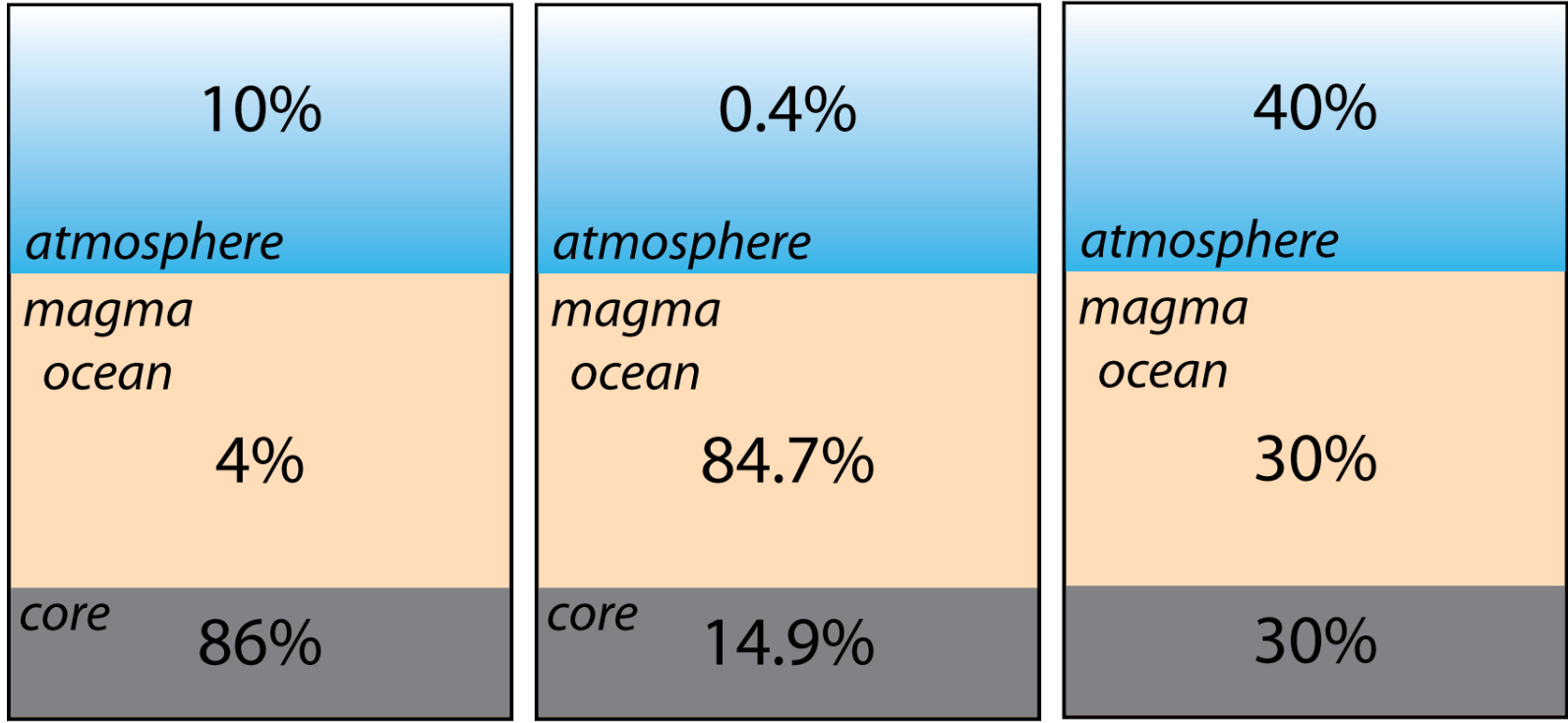
mantle ← → atmosphere



carbon

hydrogen

nitrogen

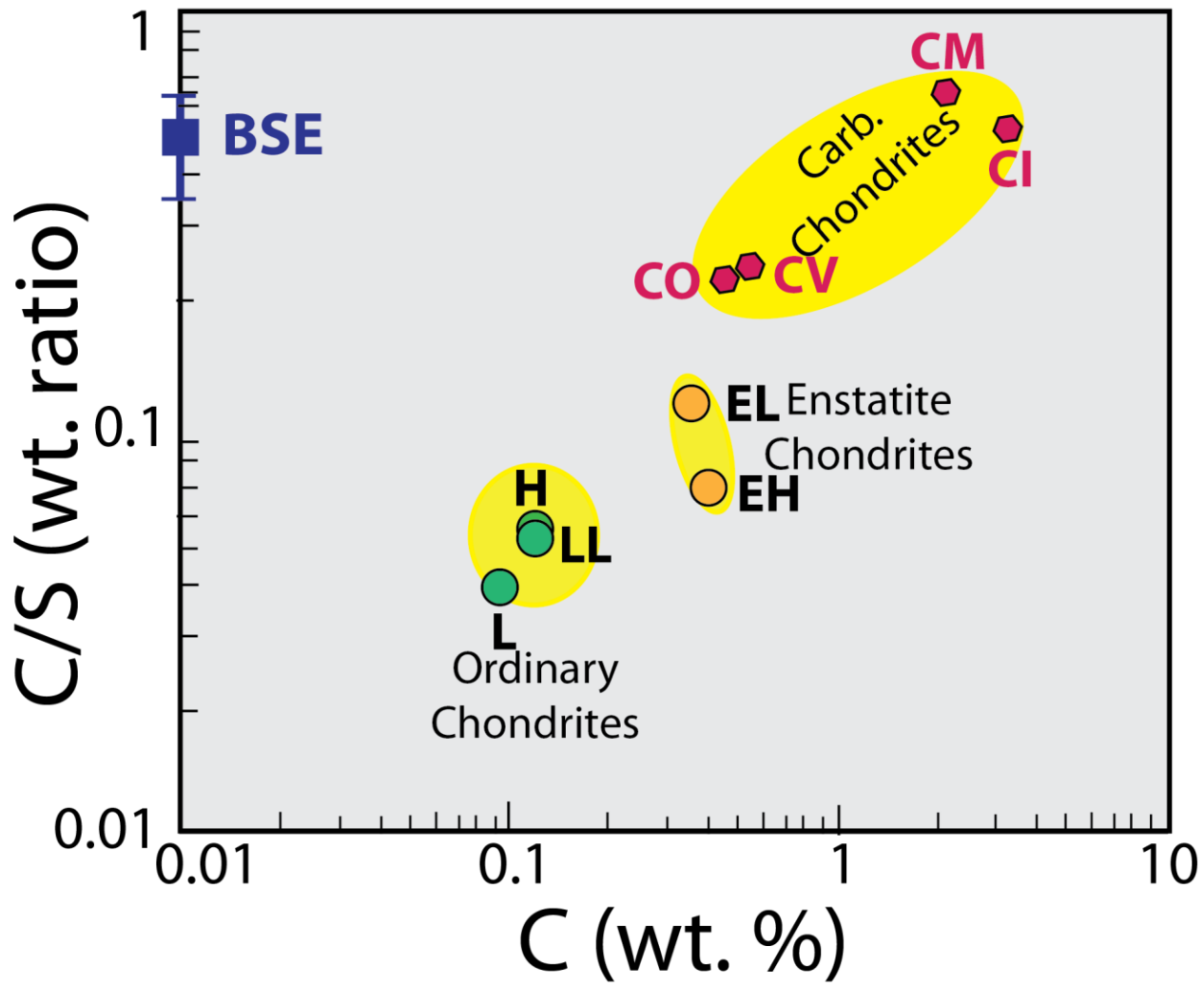


$D_C^{\text{metal/silicate}} = 200$

$D_H^{\text{metal/silicate}} = 14$

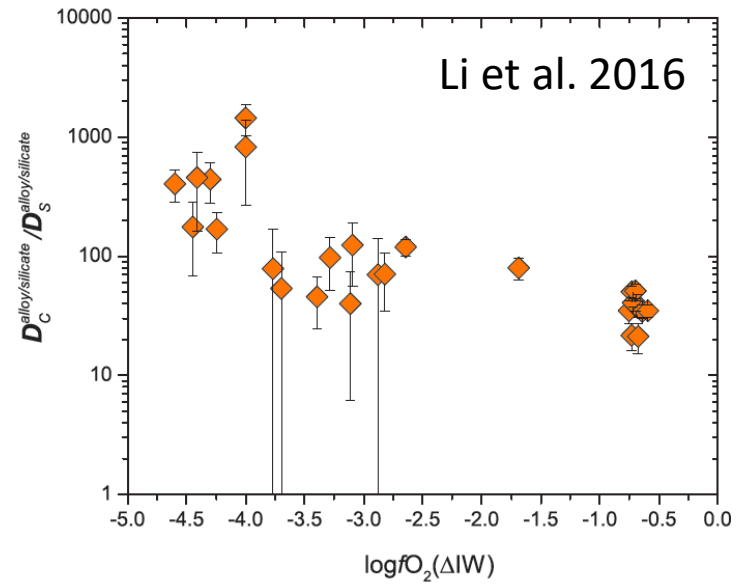
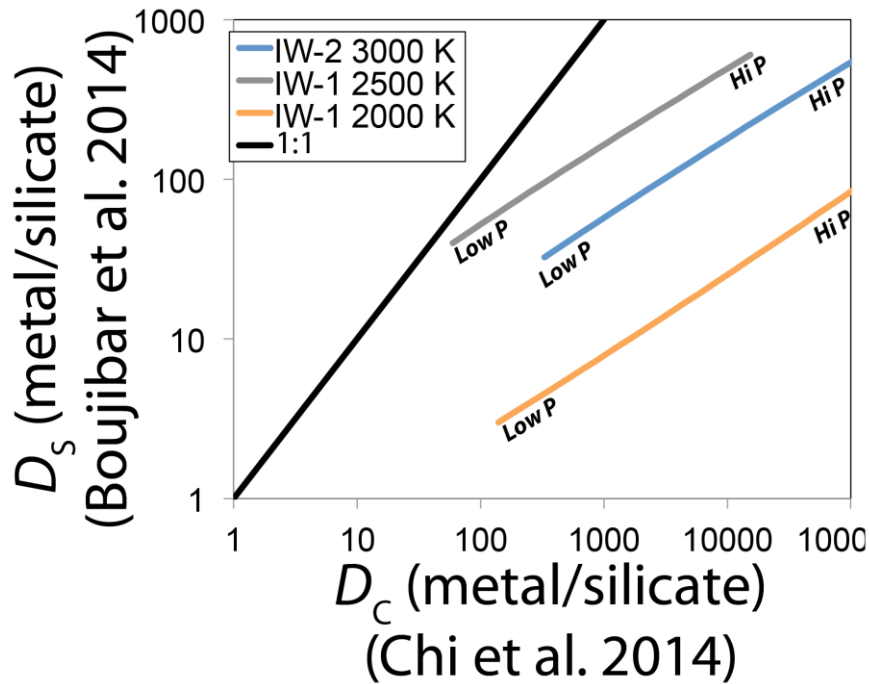
$D_N^{\text{metal/silicate}} = 10$

*Partial Mantle/Core Equilibration
(mass metal/mass mantle=0.1)*

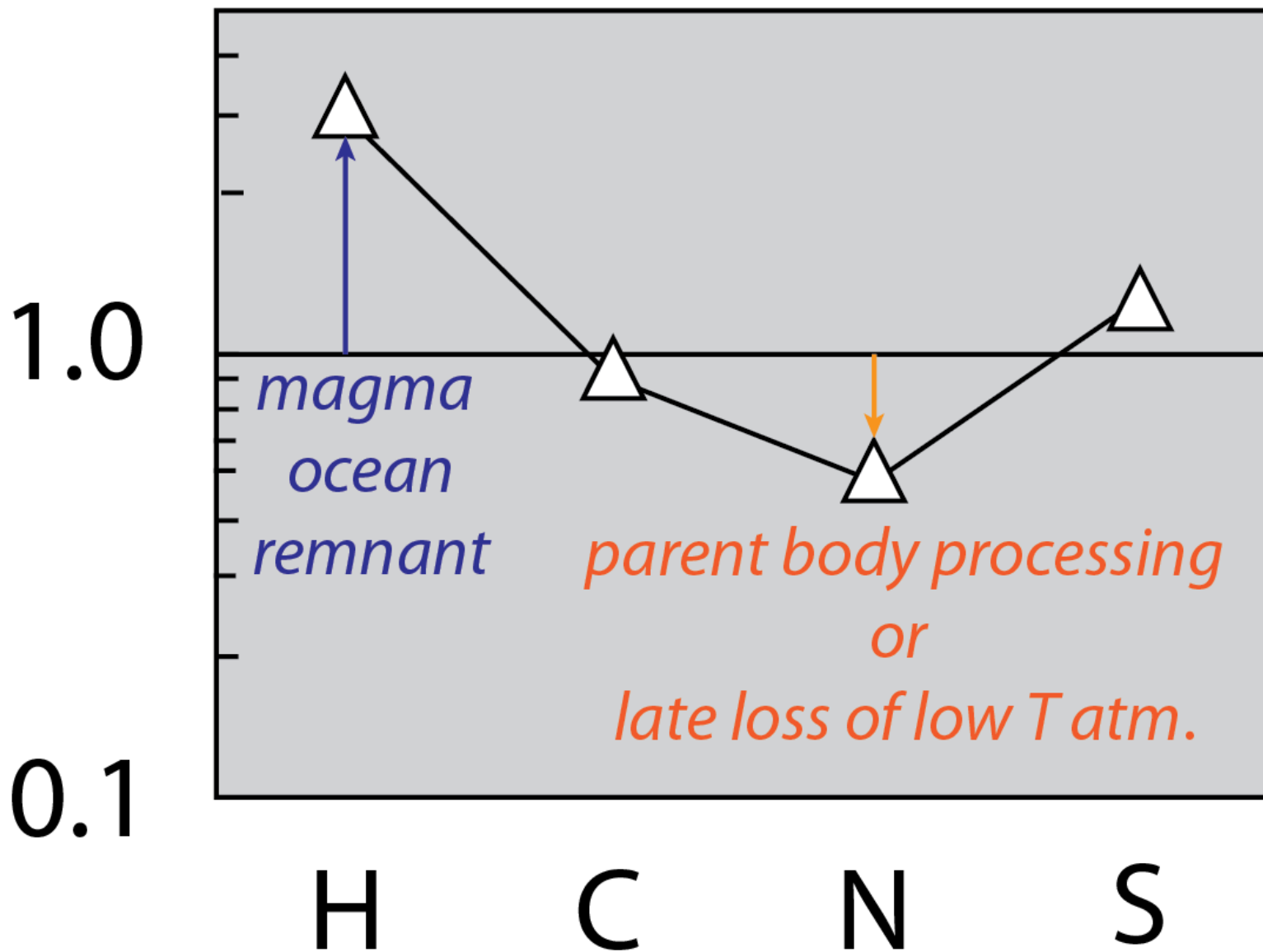


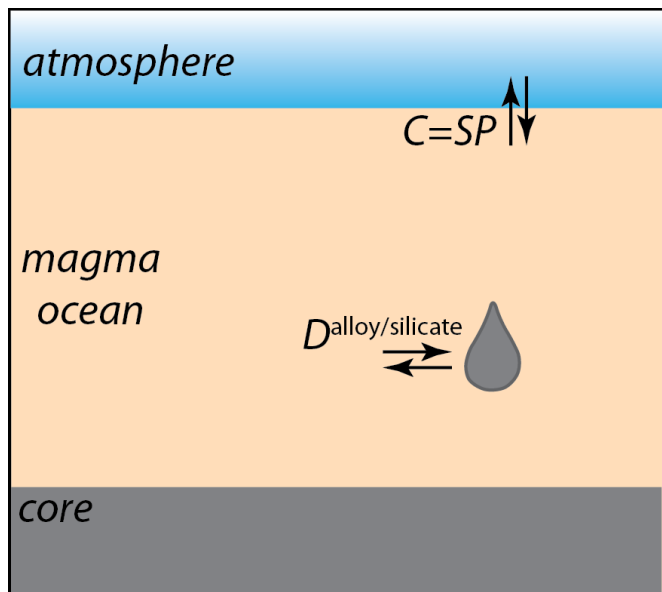
Siderophile Tendency (Preference for the core)

C>>S



BSE/Late Veneer (CI)

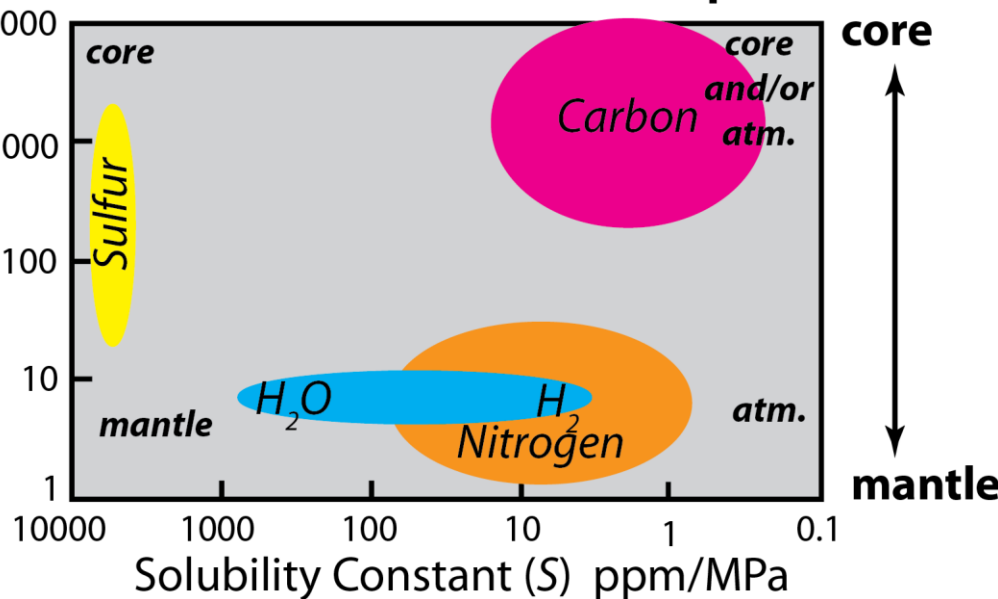




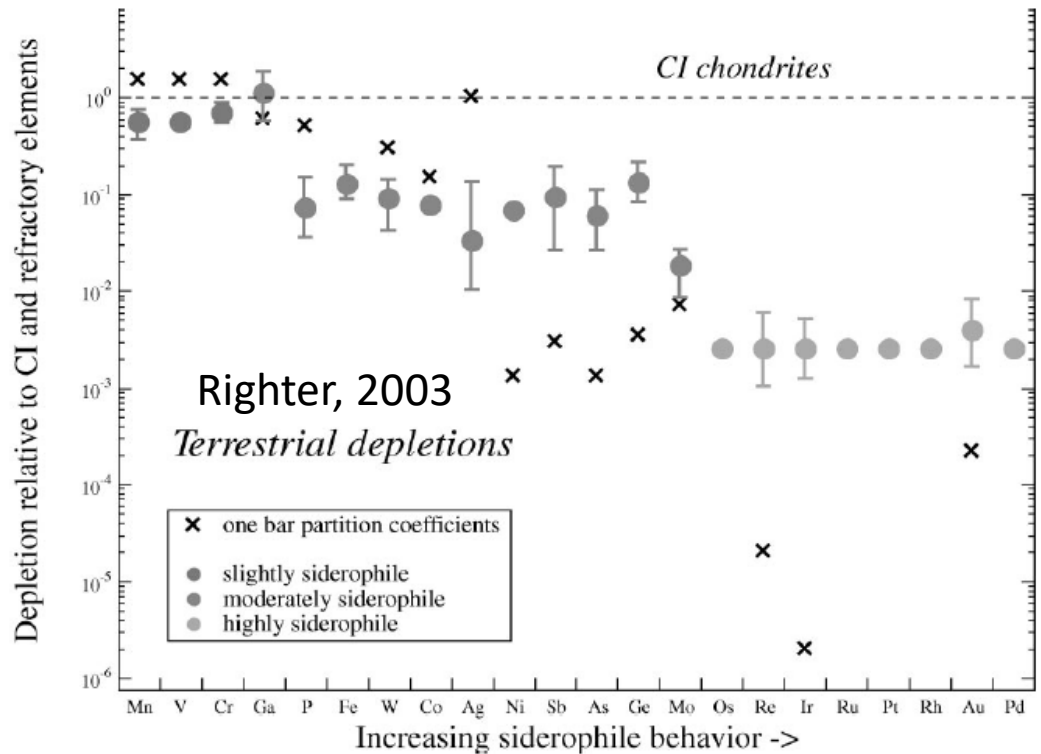
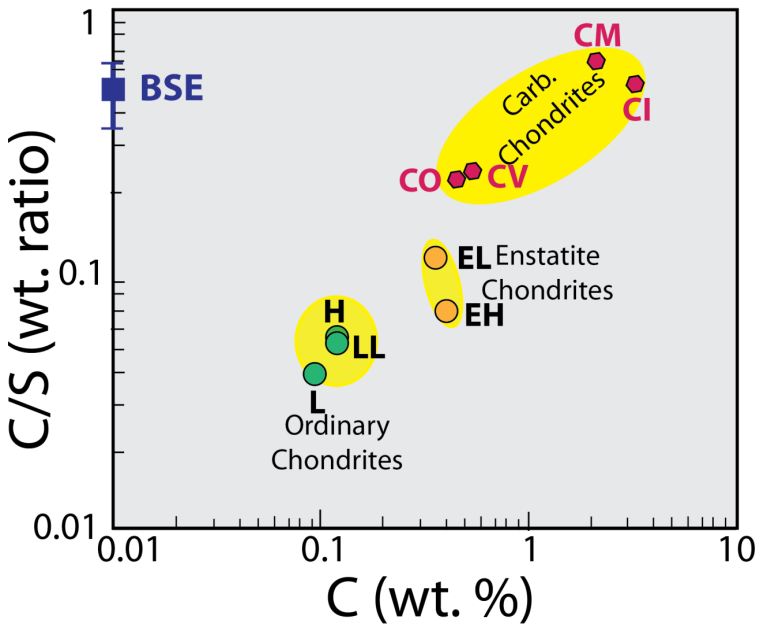
Metal/Silicate Partition

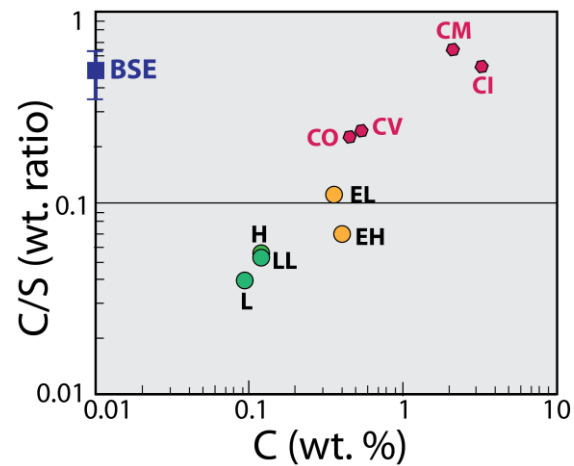
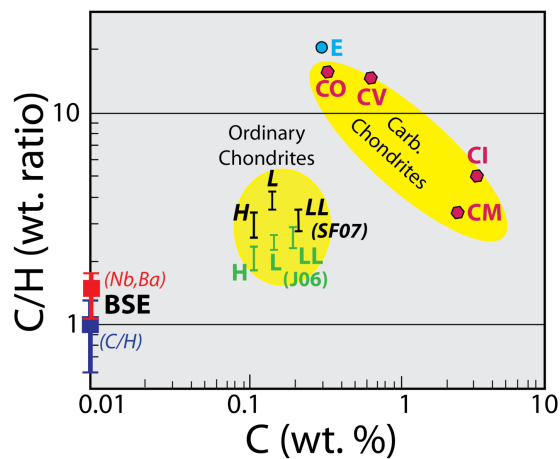
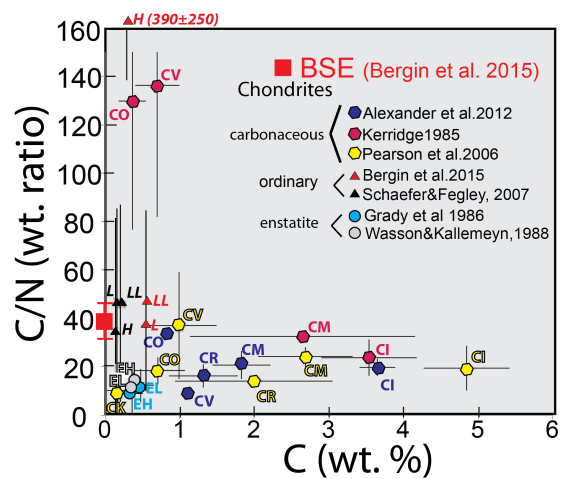
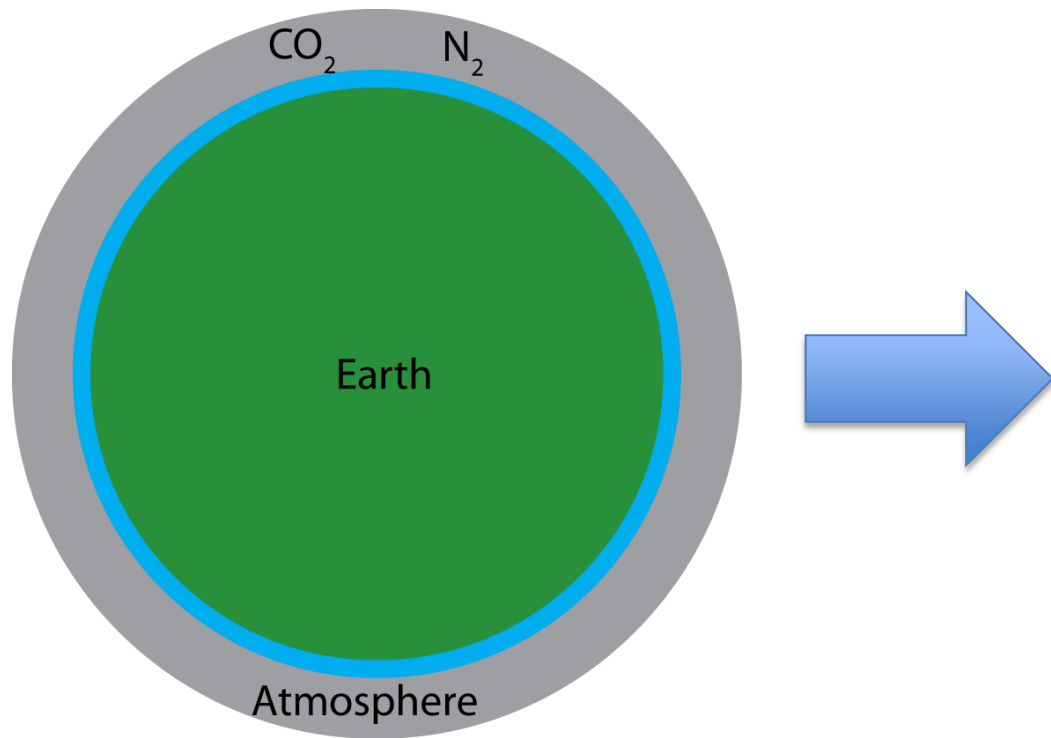
Coefficient $D_{\text{met/sil}}$

Chief Reservoirs in a Largely Molten Planet



C/S ratio has echoes of the “Late Veneer”





Summary

- The BSE budgets of C and H are well-constrained by ratios to refractory lithophile elements (Ba, Ce)
- BSE volatiles are low, but there are plenty of loss mechanisms (to core, to space)
- BSE volatiles are fractionated relative to chondrites. Either volatiles were delivered by differentiated bodies or there were selective loss mechanisms during differentiation. Probably both.