#### Dark Matter - Phonon Scattering

#### @ University of Michigan 10/09/2019



Simon Knapen Institute for Advanced Study

### There is missing mass in galaxies



#### There is missing mass in galaxies



#### Dark matter velocity

$$\langle v \rangle \sim \sqrt{\frac{G_N M_{halo}}{R_{halo}}} \sim 200 \text{ km/s}$$

Dark Matter halo

### The dark matter landscape



## The dark matter landscape



etc (ADMX, MADMAX...)

## The dark matter landscape



etc (ADMX, MADMAX...)

### **Experimental status**



Cosmic visions report 2017: 1707.04591

### Low mass dark matter detection



What do we need?

Experiment:

1. Low target mass materials:

$$q < 2m_{\chi}v_{\chi}, \qquad v_{\chi} \approx 10^{-3}$$
$$E_R = \frac{q^2}{2m_N} < 10^{-6} \times \frac{m_{\chi}^2}{m_N}$$

2. Ultra-sensitive calorimeters with low dark counts

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Theory:

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- 2. Beyond "billiard ball" scattering: structure effects are critical!

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## Experiments under development



W. Guo, D. McKinsey: 1302.0534



Example: light scalar mediator with coupling to hadrons





Anomalous cooling of stars







Anomalous cooling of stars





Dark Matter self-interactions (Relaxed for subcomponent DM)





Strong astrophysical & terrestrial constraints

SK, T. Lin, K. Zurek: 1709.07882





Experimentally viable if subcomponent DM

SK, T. Lin, K. Zurek: 1709.07882

Coupling to electrons:



 $m_{\phi} = 10^{-3} \ \mu_{\chi e}$  $10^{-36}$ SENSEI  $10^{-39}$ DAMIC SuperCDMS G2  $10^{-42}$ Dirac materials  $\overline{\sigma}_e \, [\mathrm{cm}^2]$ AI SC  $10^{-45}$  $10^{-48}$  $10^{-51}$  $\frac{\Omega_{\chi}}{\Omega_{\rm DM}}$ = 1  $10^{-54}$  $10^{-3}$  $10^{-2}$  $10^{-1}$  $10^{0}$  $10^{2}$ 10<sup>1</sup>  $10^{3}$  $m_{\chi}$  [MeV]



More freedom for mX > 1 MeV

## An important special case

Very light dark photon mediator:



#### Mediator decouples from SM at low mass No fifth force bounds (screening)

H. An, M. Pospelov, Josef Pradler, A.Ritz:1412.8378 J. Chang, R. Essig, S. McDermott: 1611.03864

. . .



. . .

 $\chi$ 

 $p^+, e^-$ 

Xenon10

 $\chi$ 

 $p^+, e^-$ 

BBN

Stellar bounds

 $10^{-32}$ 

 $10^{-33}$ 

 $10^{-34}$ 

 $\overline{\sigma}_e \,\, [\mathrm{cm}^2]$ 

R. Essig et. al.: 1509.01598

Attempt to "match" target mass with dark matter mass

- $m_{\chi} > 1 \text{ GeV}$  $\rightarrow$  nuclear recoils
- 1 MeV <  $m_{\chi}$  < 1 GeV  $\rightarrow$  electron recoils
- m<sub>{\chi}</sub> < 1 MeV

 $\rightarrow$  q  $\approx$  m<sub> $\chi$ </sub> v<sub> $\chi$ </sub> < keV  $\sim$  nm<sup>-1</sup>



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Scatter directly off phonons

→ strong material dependence!



Periodic potential (Hooke's law)



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$$\mathcal{V} = \mathcal{V}^{(0)} + \sum_{l,j} \mathcal{V}^{(i)}_{l,j} \cdot \mathbf{u}_{j,l} + \frac{1}{2} \sum_{l,l',j,j'} \mathbf{u}_{j,l} \cdot \mathcal{V}^{(2)}_{l,j,l',j'} \cdot \mathbf{u}_{j',l'} + \cdots$$
Atoms in Lattice unit cell sites





$$\mathbf{u}_{j,\mathbf{l}}(t) = \sum_{\nu}^{3n} \sum_{\mathbf{q}} \sqrt{\frac{1}{2Nm_j\omega_{\nu,\mathbf{q}}}} \left( \mathbf{e}_{\nu,j,\mathbf{q}} \hat{a}_{\nu,\mathbf{q}} e^{i\mathbf{q}\cdot(\mathbf{l}+\mathbf{r}_j^0) - i\omega_{\nu,\mathbf{q}}t} + \mathbf{e}_{\nu,j,\mathbf{q}}^* \hat{a}_{\nu,\mathbf{q}}^\dagger e^{-i\mathbf{q}\cdot(\mathbf{l}+\mathbf{r}_j^0) + i\omega_{\nu,\mathbf{q}}t} \right)$$

- phonon branches  $\nu$  :
- momentum over Brioullin zone **q** :
- atom in primitive cell j:

In the harmonic approximation, just quantize as harmonic oscillator

atomic distance



In the harmonic approximation, just quantize as harmonic oscillator

## Types of phonons

GaAs Brillouin zone





## Types of phonons

#### **Optical phonons**











DM-phonon coupling depends strongly on underlying UV physics



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## **Examples of Polar Materials**



2 atoms in primitive cell

10 atoms in primitive cell

At least two different atoms in the unit cell

## Why polar materials?

1. Optical phonons for kinematic matching



2. Natural dipole in unit cell (nanocharged DM sources tiny electric field)



SK, T. Lin, M. Pyle, K. Zurek: 1712.06598

## Why polar materials?

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2. Natural dipole in unit cell (nanocharged DM sources tiny electric field)



- 3. Semi-conductors or insulators: screening is small
- 4. Crystal axis allows for directional detection (daily modulation!)
- 5. Readily available now
## Frölich Hamiltonian

Electric dipole interacting with test charge:

H. Frölich, 1954C. Verdi, F. Giustino, Phys. Rev. Lett. 115, 176401 (2015)

$$H \sim i \, e \sum_{\mathbf{q}} \frac{\mathbf{q} \cdot \mathbf{P}}{|\mathbf{q}|^2} e^{i \mathbf{q} \cdot \mathbf{r}}$$

## Frölich Hamiltonian

H. Frölich, 1954 C. Verdi, F. Giustino, Phys. Rev. Lett. 115, 176401 (2015)



• G reciprocal lattice

## Frölich Hamiltonian

H. Frölich, 1954 C. Verdi, F. Giustino, Phys. Rev. Lett. 115, 176401 (2015)



#### Reach

Both GaAs and Sapphire probe Dark Matter masses as low as 10 keV



Probe the new parameter space with milligram-day exposure

SK, T. Lin, M. Pyle, K. Zurek: 1712.06598 S. Griffin, SK, T. Lin, M. Pyle, K. Zurek: 1807.10291

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 $\chi$ 

 $p^+, e^-$ 

 $\chi$ 

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19

 $\chi$ 

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 $\chi$ 













#### Amplitude and pattern depends on DM mass

# Daily modulation (dark photon mediator)

 $p^+, e^-$ 

 $p^+, e$ 



## Daily modulation (dark photon mediator)

Amplitude and pattern depends on DM mass

22

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#### Amplitude and pattern depends on DM mass

#### Dark photon absorption

Very light, bosonic dark matter can be absorbed on the target

Example: Dark photon dark matter:  $\mathcal{L} \supset -\frac{\kappa}{2} F'_{\mu\nu} F^{\mu\nu}$ 

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## Dark Matter coupling to phonons



Which modes to use?





1. Single acoustic:

Experimentally extremely challenging



 $\chi$ 

N

 $\chi$ 

N

 $\phi$ 

1. Single acoustic:

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2. Single optical:

Strong destructive interference



 $\chi$ 

N

 $\chi$ 

N

 $\phi$ 

1. Single acoustic:

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3. Double acoustic:

Next-to-leading order







 $\chi$ 

N

 $\chi$ 

N

 $\phi$ 

Scattering potential:

$$V(\mathbf{r}) = \frac{2\pi b_X}{m_X} \sum_{\ell,j} A_j \delta(\mathbf{r}_{\ell,j} - \mathbf{r})$$

Phonon form factor:

$$|F_{\nu}(\mathbf{q})|^{2} = \left|\sum_{j} \frac{A_{j}}{\sqrt{m_{j}}} e^{-W_{j}(\mathbf{q})} \mathbf{q} \cdot \mathbf{e}_{\nu,j,\mathbf{q}} e^{-i\mathbf{q} \cdot \mathbf{r}_{j}}\right|^{2}$$



SK, T. Lin, M. Pyle, K. Zurek: 1712.06598S. Griffin, SK, T. Lin, M. Pyle, K. Zurek: 1807.10291P. Cox, T. Melia, S. Rajendran: 1905.05575

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Daily modulation

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Acoustic
$$|F_{\nu}^{(ac)}(\mathbf{q})|^{2} \approx \frac{q^{2}}{m_{p}} \left|\sum_{j} \sqrt{A_{j}} e^{-i\mathbf{q}\cdot\mathbf{r}_{j}}\right|^{2}$$

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Phonon form factor:





Lattice constant

Destructive interference kills leading order piece for optical phonons

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Daily modulation

## Rate for scalar mediator



Huge enhancement for the acoustic mode (but needs ultra low threshold)

SK, T. Lin, M. Pyle, K. Zurek: 1712.06598 S. Griffin, SK, T. Lin, M. Pyle, K. Zurek: 1807.10291

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 $\chi$ 

N

 $\chi$ 

N

 $\phi$ 

# Daily modulation (Sapphire)





Qualitatively different from dark photon mediator!

S. Griffin, SK, T. Lin, M. Pyle, K. Zurek: 1807.10291

N

#### Which modes to use?

1. Single acoustic:

Experimentally extremely challenging

2. Single optical:

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3. Double acoustic:

Next-to-leading order







In progress with B. Campbell-Deem, T. Lin, P. Cox, T. Melia

 $\chi$ 

N

 $\chi$ 

N

 $\phi$ 

## Phonon perturbation theory

Scattering potential:

$$V(\mathbf{r}) = \frac{2\pi b_X}{m_X} \sum_{\ell,j} A_j \delta(\mathbf{r}_{\ell,j} - \mathbf{r}) \qquad \Rightarrow \qquad V(\mathbf{q}) = \frac{2\pi b_X}{m_X} \sum_{\ell,j} A_j e^{i\mathbf{q}\cdot\mathbf{r}_{\ell,j}}$$

Matrix element:

$$\left|\mathcal{M}\right|^{2} \sim \left|\left\langle \Phi_{f}\right| \sum_{\ell,l} A_{j} e^{i\mathbf{q}\cdot\mathbf{r}_{\ell,j}} \left|0\right\rangle\right|^{2}$$

Double expansion in the phonon self-coupling and momentum transfer

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Double expansion in the phonon self-coupling and momentum transfer

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 $\phi$ 

## **Phonon self-interactions**

Potential:

$$\mathcal{V} = \mathcal{V}^{(0)} + \sum_{\mathbf{l},j} \mathcal{V}^{(i)}_{\mathbf{l},j} \cdot \mathbf{u}_{j,\mathbf{l}} + \frac{1}{2} \sum_{\mathbf{l},\mathbf{l}',j,j'} \mathbf{u}_{j,\mathbf{l}} \cdot \mathcal{V}^{(2)}_{\mathbf{l},j,\mathbf{l}',j'} \cdot \mathbf{u}_{j',\mathbf{l}'} + \cdots$$



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Higher order terms give effective Hamiltonian (in isotropic approximation):

$$\delta H = \int d^3 \mathbf{r} \; \frac{1}{2} (\beta + \lambda) u_{ii} u_{jk} u_{jk} + (\gamma + \mu) u_{ij} u_{ki} u_{kj} + \frac{\alpha}{3!} u_{ii} u_{jj} u_{kk} + \frac{\beta}{2} u_{ii} u_{jk} u_{kj} + \frac{\gamma}{3} u_{ij} u_{jk} u_{ki} u_{ki} + \frac{\beta}{3!} u_{ii} u_{jk} u_{ki} + \frac{\beta}{3!} u_{ii} u_{ii} u_{jk} u_{ki} + \frac{\beta}{3!} u_{ii} u_{jk} u_{ki} u_{ki} + \frac{\beta}{3!} u_{ii} u_{jk} u_{ki} u_{ki}$$

With  $u_{ij} \equiv \partial_i u_j$ 

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With  $u_{ij} \equiv \partial_i u_j$ 

Couplings arise from expanding the potential beyond harmonic approximation



Third order elastic constants

Can be measured!

## Self-interactions matrix element

$$\begin{split} \widetilde{\mathcal{M}} &= (\beta + \lambda) \Big[ (\mathbf{q} \cdot \mathbf{e}) (\mathbf{k}_1 \cdot \mathbf{k}_2) (\mathbf{e}_1 \cdot \mathbf{e}_2) + (\mathbf{k}_1 \cdot \mathbf{e}_1) (\mathbf{q} \cdot \mathbf{k}_2) (\mathbf{e} \cdot \mathbf{e}_2) + (\mathbf{k}_2 \cdot \mathbf{e}_2) (\mathbf{k}_1 \cdot \mathbf{q}) (\mathbf{e}_1 \cdot \mathbf{e}) \Big] \\ &+ (\gamma + \mu) \Big[ (\mathbf{q} \cdot \mathbf{k}_2) \big[ (\mathbf{k}_2 \cdot \mathbf{e}_1) (\mathbf{e}_2 \cdot \mathbf{e}) + (\mathbf{k}_2 \cdot \mathbf{e}) (\mathbf{e}_2 \cdot \mathbf{e}_1) \big] \\ &+ (\mathbf{k}_2 \cdot \mathbf{k}_1) \big[ (\mathbf{q} \cdot \mathbf{e}_1) (\mathbf{e}_2 \cdot \mathbf{e}) + (\mathbf{q} \cdot \mathbf{e}_2) (\mathbf{e} \cdot \mathbf{e}_1) \big] \\ &+ (\mathbf{q} \cdot \mathbf{k}_2) \big[ (\mathbf{k}_1 \cdot \mathbf{e}_2) (\mathbf{e}_1 \cdot \mathbf{e}) + (\mathbf{k}_1 \cdot \mathbf{e}) (\mathbf{e}_1 \cdot \mathbf{e}_2) \big] \Big] \\ &+ \alpha (\mathbf{q} \cdot \mathbf{e}) (\mathbf{k}_1 \cdot \mathbf{e}_1) (\mathbf{k}_2 \cdot \mathbf{e}_2) \\ &+ \beta \Big[ (\mathbf{k}_1 \cdot \mathbf{e}_1) (\mathbf{q} \cdot \mathbf{e}_2) (\mathbf{k}_2 \cdot \mathbf{e}) + (\mathbf{q} \cdot \mathbf{e}) (\mathbf{k}_1 \cdot \mathbf{e}_2) (\mathbf{k}_2 \cdot \mathbf{e}_1) + (\mathbf{k}_2 \cdot \mathbf{e}_2) (\mathbf{q} \cdot \mathbf{e}_1) (\mathbf{k}_1 \cdot \mathbf{e}) \Big] \\ &+ \gamma \Big[ (\mathbf{q} \cdot \mathbf{e}_1) (\mathbf{k}_1 \cdot \mathbf{e}_2) (\mathbf{k}_2 \cdot \mathbf{e}) + (\mathbf{q} \cdot \mathbf{e}_1) (\mathbf{k}_1 \cdot \mathbf{e}) (\mathbf{k}_2 \cdot \mathbf{e}_1) \Big] \end{split}$$



In progress with B. Campbell-Deem, T. Lin, P. Cox, T. Melia

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4 non-vanishing channels:

- $LA^* \rightarrow LA LA$
- LA\*  $\rightarrow$  LA TA
- LA\*  $\rightarrow$  TA TA (polarized in momentum-plane)
- LA\*  $\rightarrow$  TA TA (polarized orthogonal to momentum-plane)



# Comparing different channels

#### Example GaAs:

Differential scattering rate


# **Comparing different channels**

#### Example GaAs:

#### Differential scattering rate



The anharmonic contribution tends to dominate

### **Preliminary results**









In progress with B. Campbell-Deem, T. Lin, P. Cox, T. Melia

#### 

Translations spontaneously broken, so proportional to momenta

# $\begin{array}{l} \underline{\operatorname{Crystal}} \\ \langle \mathbf{q} \,|\, \delta H \,|\, \mathbf{k}_1, \, \mathbf{k}_2 \rangle \sim (\mathbf{q} \cdot \mathbf{e}) \,(\mathbf{k}_1 \cdot \mathbf{e}_1) \,(\mathbf{k}_2 \cdot \mathbf{e}_2) + \cdots \\ \sim q \end{array} \longrightarrow \qquad \begin{array}{l} \operatorname{Rate} \ \sim \mathcal{O}(q^2) \\ \end{array}$

Translations spontaneously broken, so proportional to momenta

#### Superfluid Helium

$$\langle \mathbf{q} \, | \, \delta H \, | \, \mathbf{k}_1, \, \mathbf{k}_2 \rangle \sim \lambda \mathbf{q} \cdot (\mathbf{k}_1 + \mathbf{k}_2) + \lambda' \omega \, \mathbf{k}_1 \cdot \mathbf{k}_2 + \lambda'' \omega \omega_1 \omega_2 + \cdots$$
$$\sim q^2$$

Translations NOT spontaneously broken, but mysterious cancellation anyways...

#### 

#### Superfluid Helium



Translations NOT spontaneously broken, but mysterious cancellation anyways...

~

# $\begin{array}{l} \underbrace{\operatorname{Crystal}}_{\langle \mathbf{q} | \,\delta H \, | \, \mathbf{k}_{1}, \, \mathbf{k}_{2} \rangle \sim (\mathbf{q} \cdot \mathbf{e}) \, (\mathbf{k}_{1} \cdot \mathbf{e}_{1}) \, (\mathbf{k}_{2} \cdot \mathbf{e}_{2}) + \cdots}_{\sim q} & \longrightarrow & \operatorname{Rate} \ \sim \mathcal{O}(q^{2}) \\ \\ \text{Translations spontaneously broken, so proportional to momenta} \\ \\ \underbrace{\operatorname{Superfluid Helium}}_{\langle \mathbf{q} | \,\delta H \, | \, \mathbf{k}_{1}, \, \mathbf{k}_{2} \rangle \sim \lambda \mathbf{q} \cdot (\mathbf{k}_{1} + \mathbf{k}_{2}) + \lambda' \omega \mathbf{k}_{1} \cdot \mathbf{k}_{2} + \lambda'' \omega \omega_{1} \omega_{2} + \cdots} & \longrightarrow & \operatorname{Rate} \ \sim \mathcal{O}(q^{4}) \\ & \sim q^{2} \end{array}$



Translations NOT spontaneously broken, but mysterious cancellation anyways...

#### Crossing symmetry then implies that the leading term in the q-expansion cancels

(The self-interaction is however much stronger in helium, overcoming this additional suppression)

SK, T. Lin, K. Zurek: 1611.06228 A. Caputo, A. Esposito, A. Polosa: 1907.10635

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SK, T. Lin, M. Pyle, K. Zurek: 1712.06598
 S. Griffin, SK, T. Lin, M. Pyle, K. Zurek: 1807.10291
 P. Cox, T. Melia, S. Rajendran: 1905.05575
 B. Campbell-Deem, P. Cox, T. Melia, SK, T. Lin: to appear

[a] K. Schutz, K. Zurek: 1604.08206
[b] SK, T. Lin, K. Zurek: 1611.06228
[c] F. Acanfora, A. Esposito, A. Polosa: 1902.02361
[d] A. Caputo, A. Esposito, A. Polosa: 1907.10635

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Summary		TES and QP collection antennas (W) Athermal Phonon Collection Fins (A) 1 cm <sup>3</sup> Polar Crystal	N 25 M M 25 M
Scatt	erina	Crystals	Superfluid Helium
	<u> </u>		
Coupling to charge	Single optical	Large [1,2]	X
	Single acoustic	Tiny [1,2]	Tiny [b]
Coupling to mass	Single acoustic	experimentally hard [1,2]	experimentally impossible
	Single optical	small [1,2,3]	X
	multiphonon	small [4]	small [a,b,c,d]

SK, T. Lin, M. Pyle, K. Zurek: 1712.06598
 S. Griffin, SK, T. Lin, M. Pyle, K. Zurek: 1807.10291
 P. Cox, T. Melia, S. Rajendran: 1905.05575
 B. Campbell-Deem, P. Cox, T. Melia, SK, T. Lin: to appear

[a] K. Schutz, K. Zurek: 1604.08206
[b] SK, T. Lin, K. Zurek: 1611.06228
[c] F. Acanfora, A. Esposito, A. Polosa: 1902.02361
[d] A. Caputo, A. Esposito, A. Polosa: 1907.10635

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Summary		TES and QP collection antennas (W) Athermal Phonon Collection Fins (A) 1 cm <sup>3</sup> Polar Crystal	M N N N
<u>Scatt</u>	ering	Crystals	Superfluid Helium
Coupling to charge	Single optical Single acoustic	Large [1,2] Tiny [1,2]	X Tiny [b]
	Single acoustic	experimentally hard [1,2]	experimentally impossible
Coupling to mass	Single optical	small [1,2,3]	X
	multiphonon	small [4]	small [a,b,c,d]
Δ <b>Ι</b> <sub>2</sub>			

#### <u>Absorption</u>

Dark photon	Large [2]	Tiny [b]
Scalar	Future work	Future work

SK, T. Lin, M. Pyle, K. Zurek: 1712.06598
 S. Griffin, SK, T. Lin, M. Pyle, K. Zurek: 1807.10291
 P. Cox, T. Melia, S. Rajendran: 1905.05575
 B. Campbell-Deem, P. Cox, T. Melia, SK, T. Lin: to appear

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